

# Singular spectrum analysis of GRACE observations

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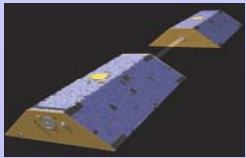
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# MSSA IN APPLICATION TO GRACE MONTHLY DATA



Gravity Recovery and Climate Experiment (GRACE) satellites, launched 17.03.2002, provides an unprecedented set of Earth's temporal gravity field observations, whose signal manifests as Earth's mass redistribution within the Earth system. However, de-stripping/filtering to diminish the correlated high frequency errors is currently required to use the GRACE spherical harmonic data products.



We applied the MSSA to monthly GRACE gravity field solutions in both the spatial and spectral domains. The principal components (PC) were identified to be related to different physical processes, including seasonal to interannual water mass redistribution, ice/glacial melting, GIA, Sumatra-earthquake. Preliminary results indicate that one could separate noise and signals using the MSSA approach.

## INITIAL DATA IN SPECTRAL DOMAIN

We used 79 CSR Level-2 RL04 monthly GRACE data since 2002 till 2009 with spherical harmonic coefficients complete to degree 60.

$$V(\rho, \theta, \lambda) = \frac{GM}{\rho} \sum_{n=0}^{\infty} \sum_{m=0}^n \left(\frac{a}{\rho}\right)^n (\bar{c}_{nm} \cos m\lambda + \bar{s}_{nm} \sin m\lambda) \bar{P}_{nm}(\cos \theta).$$

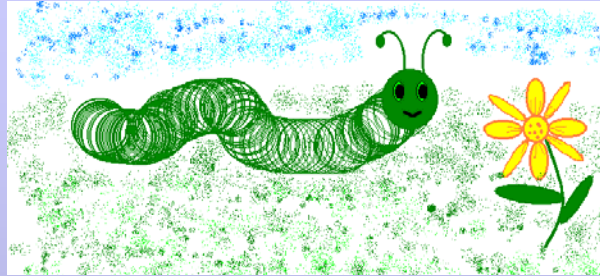
$$\bar{P}_{nm}(\cos \theta) = \sqrt{\delta_m (2n+1) \frac{(n-m)!}{(n+m)!}} \sin^m \theta \frac{d^m P_n(\cos \theta)}{d(\cos \theta)^m}$$

Degree  $n$  changes from 2 for the appropriate selection of coordinates.

$C_{20}$  coefficients were excluded in processing. GGM01C model was used as a reference and subtracted. Results are represented in terms of equivalent water height (EWH) levels (mm).

$$\Delta h(\delta, \lambda, t) = \frac{a\rho_{ave}}{3\rho_w} \sum_{n=2}^{60} \sum_{m=0}^n \frac{2n+1}{1+k_n} W_n (\Delta \bar{c}_{nm}(t) \cos m\lambda + \Delta \bar{s}_{nm}(t) \sin m\lambda) \bar{P}_{nm}(\cos \theta)$$

# “CATERPILLAR” – SSA METHOD



1

Trajectory  
matrix

$$X = \begin{bmatrix} X_1 & \dots & X_K \end{bmatrix}$$

$$[L \times K]$$

2

Singular value decomposition SVD

$$X = USV^T = \sqrt{\lambda_1} U_1 V_1^T + \dots + \sqrt{\lambda_d} U_d V_d^T$$

$$X = \begin{bmatrix} f_1 & f_2 & \dots & f_{N-L} \\ f_1 & f_2 & \dots & f_{N-L+1} \\ \vdots & \vdots & \ddots & \vdots \\ f_{L-1} & f_{L-1} & \dots & f_{N-1} \end{bmatrix}$$

3

$X = [x_1, x_2, \dots]$

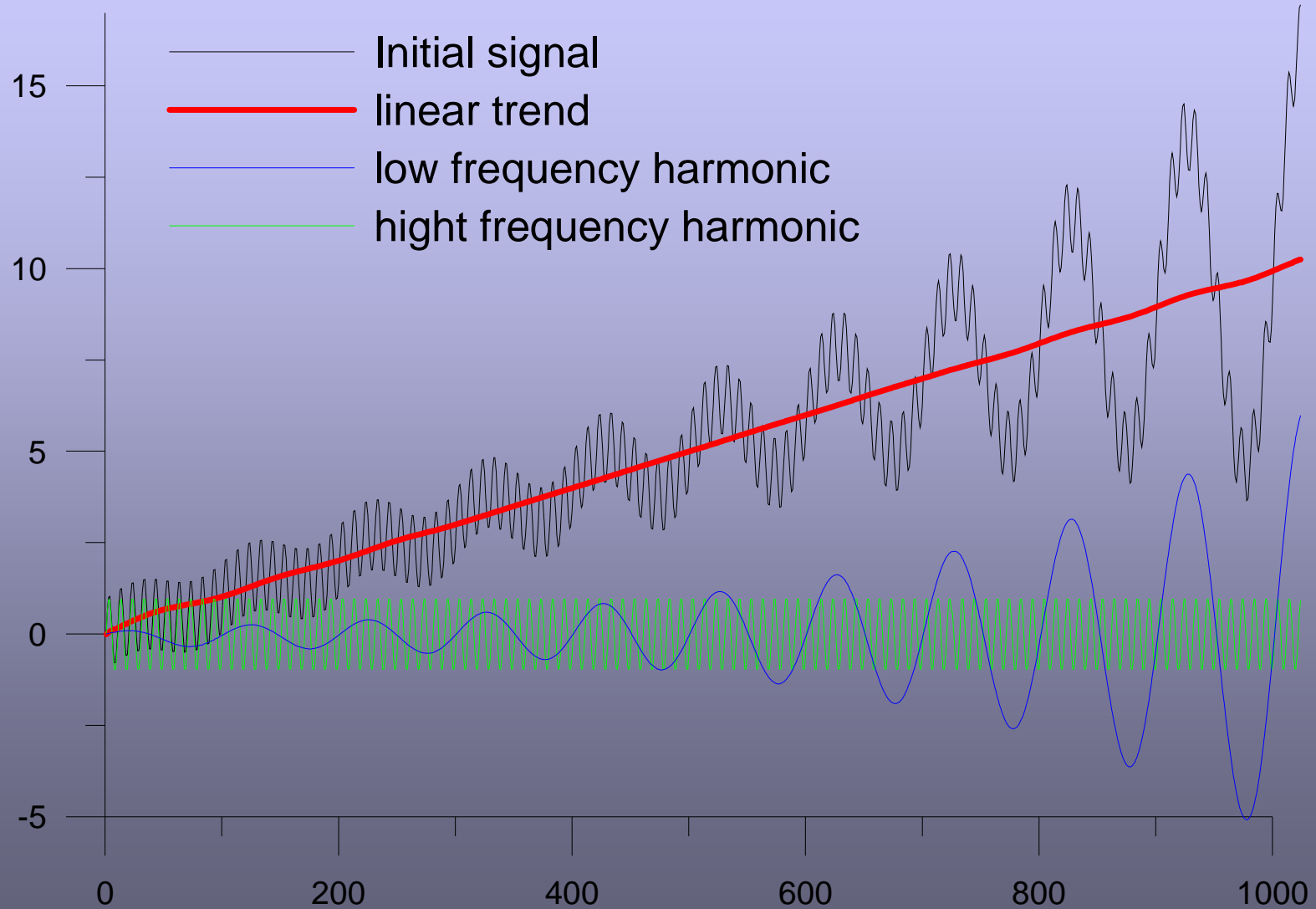
Principal components grouping

4

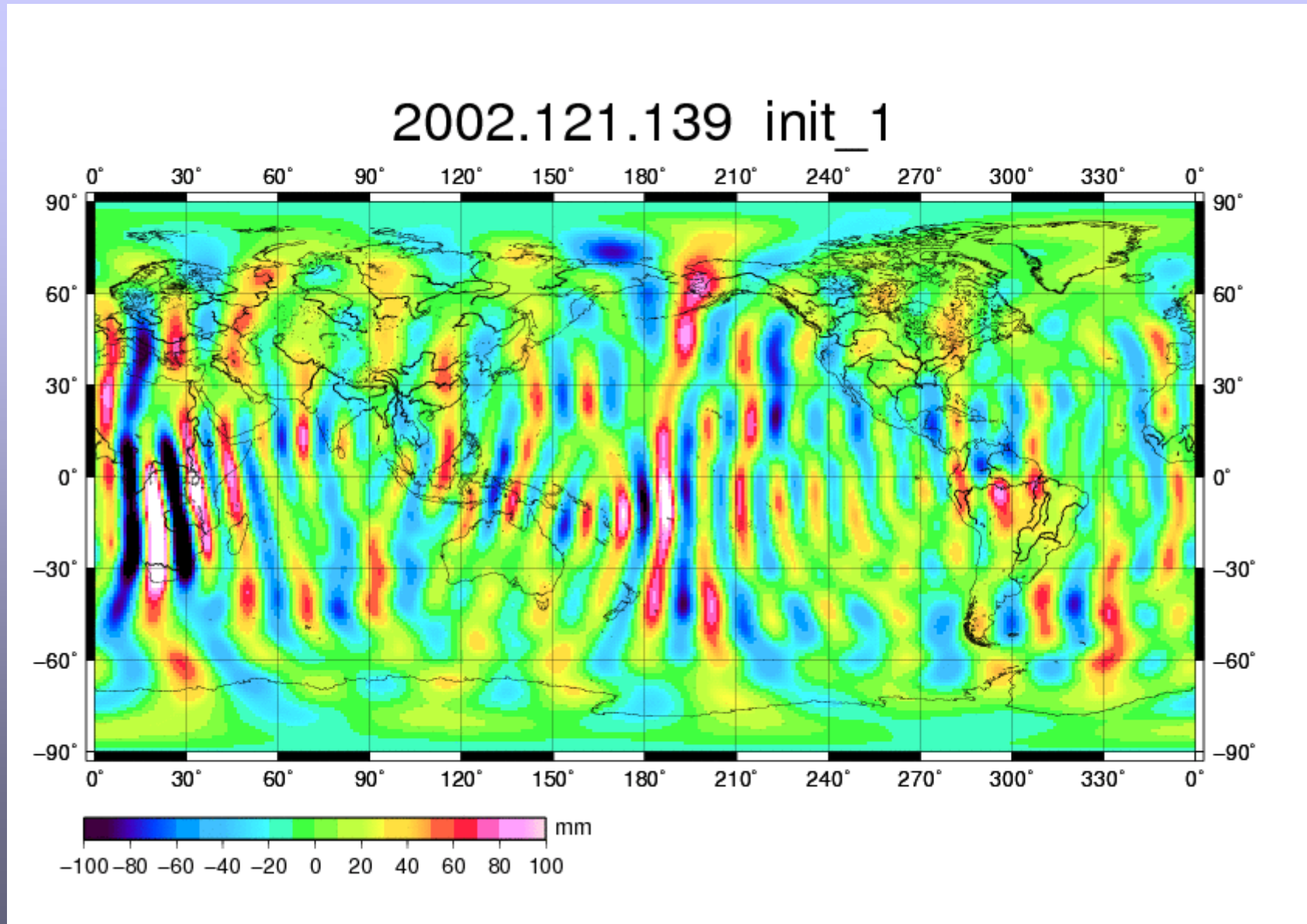
Hankelization

$$g_k = \begin{cases} \frac{1}{k+1} \sum_{m=1}^{k+1} y_{m, k-m+2}, & 0 \leq k < L^* - 1 \\ \frac{1}{L^*} \sum_{m=1}^{L^*} y_{m, k-m+2}, & L^* - 1 \leq k < K^* \\ \frac{1}{N+K} \sum_{m=k-K^*+2}^{N-K^*+1} y_{m, k-m+2}, & K^* \leq k < N \end{cases}$$

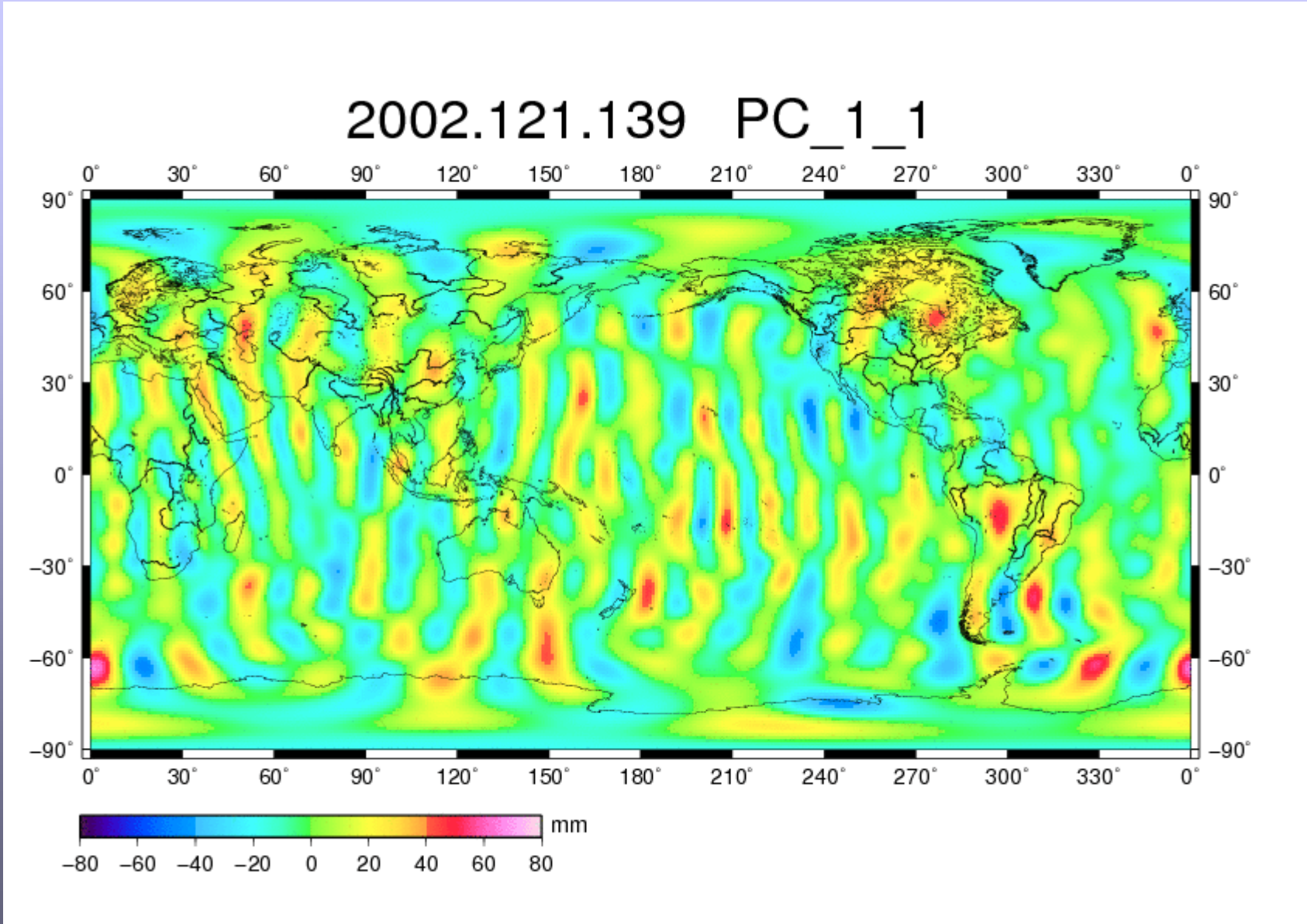
# SIMPLE EXAMPLE OF 1D SSA



# INITIAL UNFILTERED GRACE DATA (SPATIAL MAP)

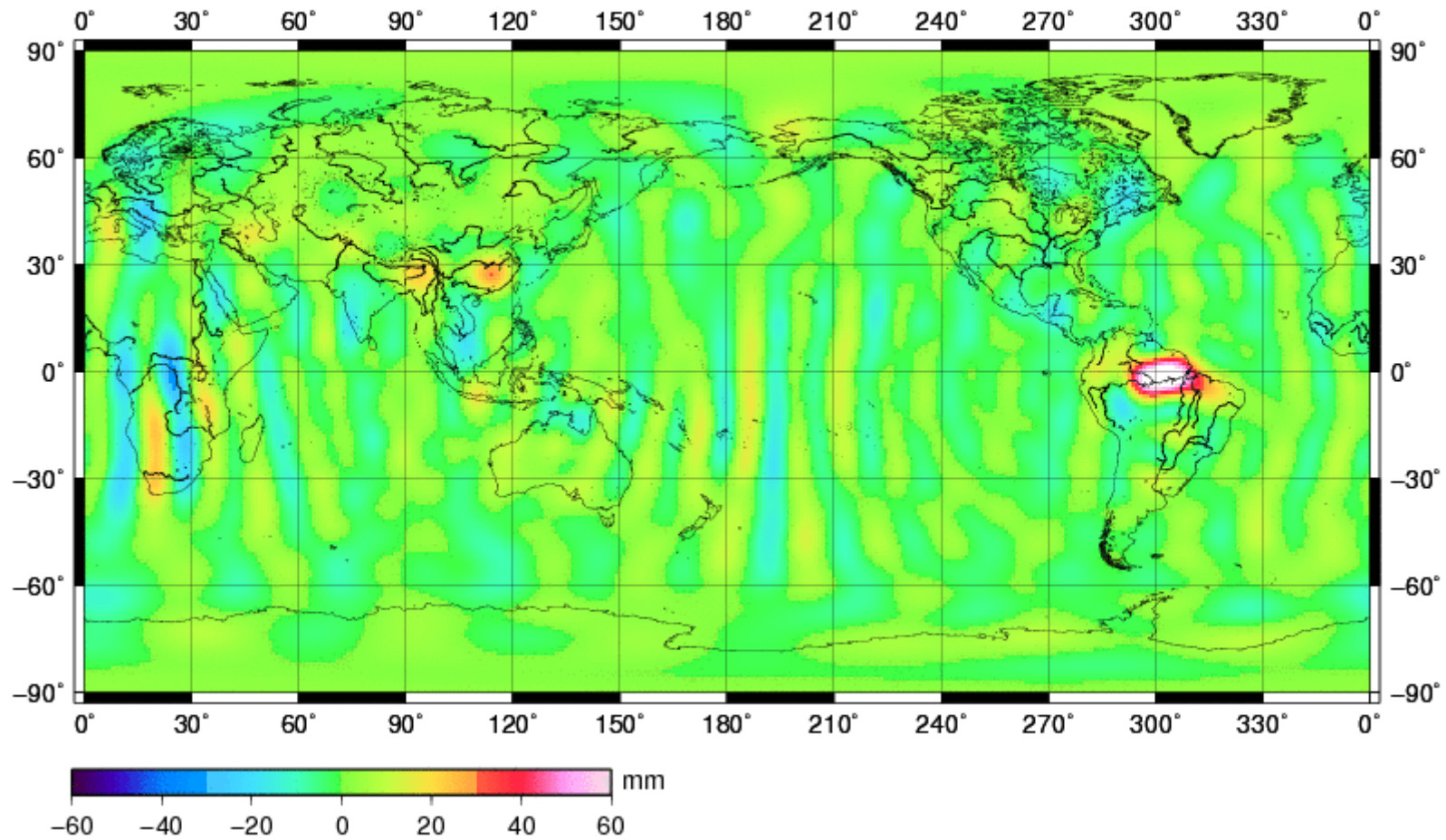


# Principal Component N1 for MSSA with L=10 in spectral domain



# PC 2+3 spectral MSSA, L=10

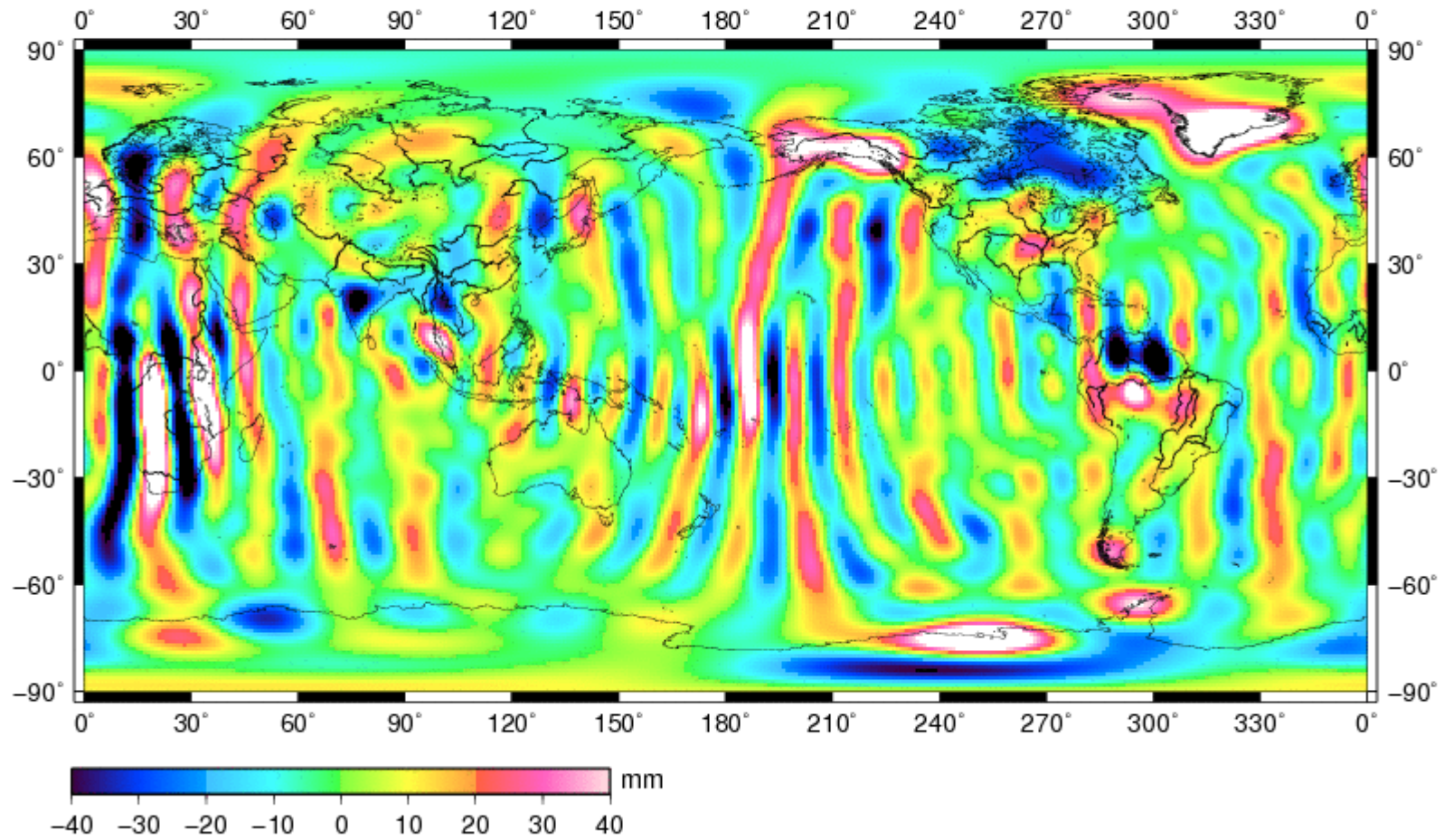
2002.121.139 PC 2+3\_1





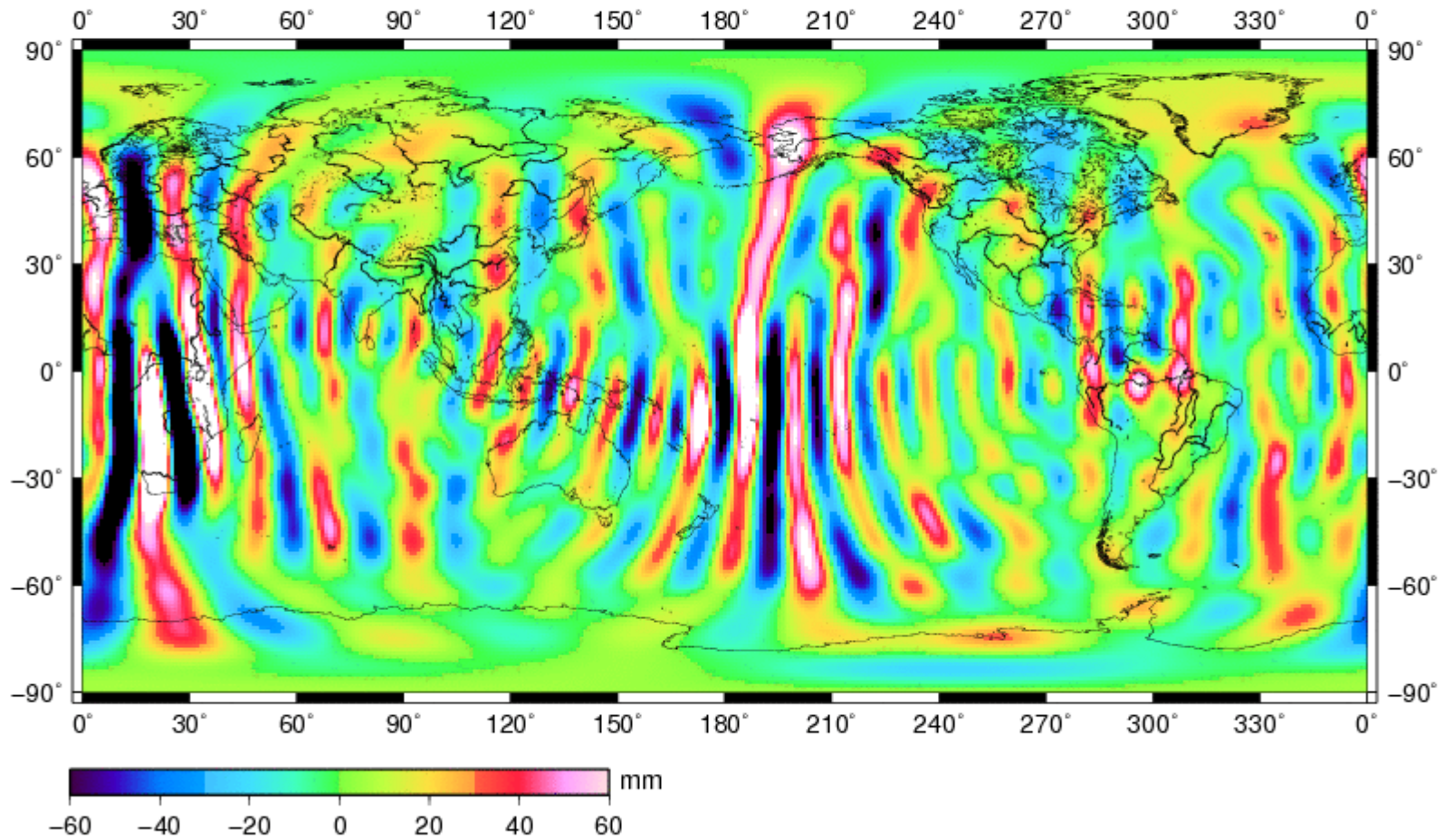
# PC4, spectral MSSA, L=10

2002.121.139 PC\_4\_1



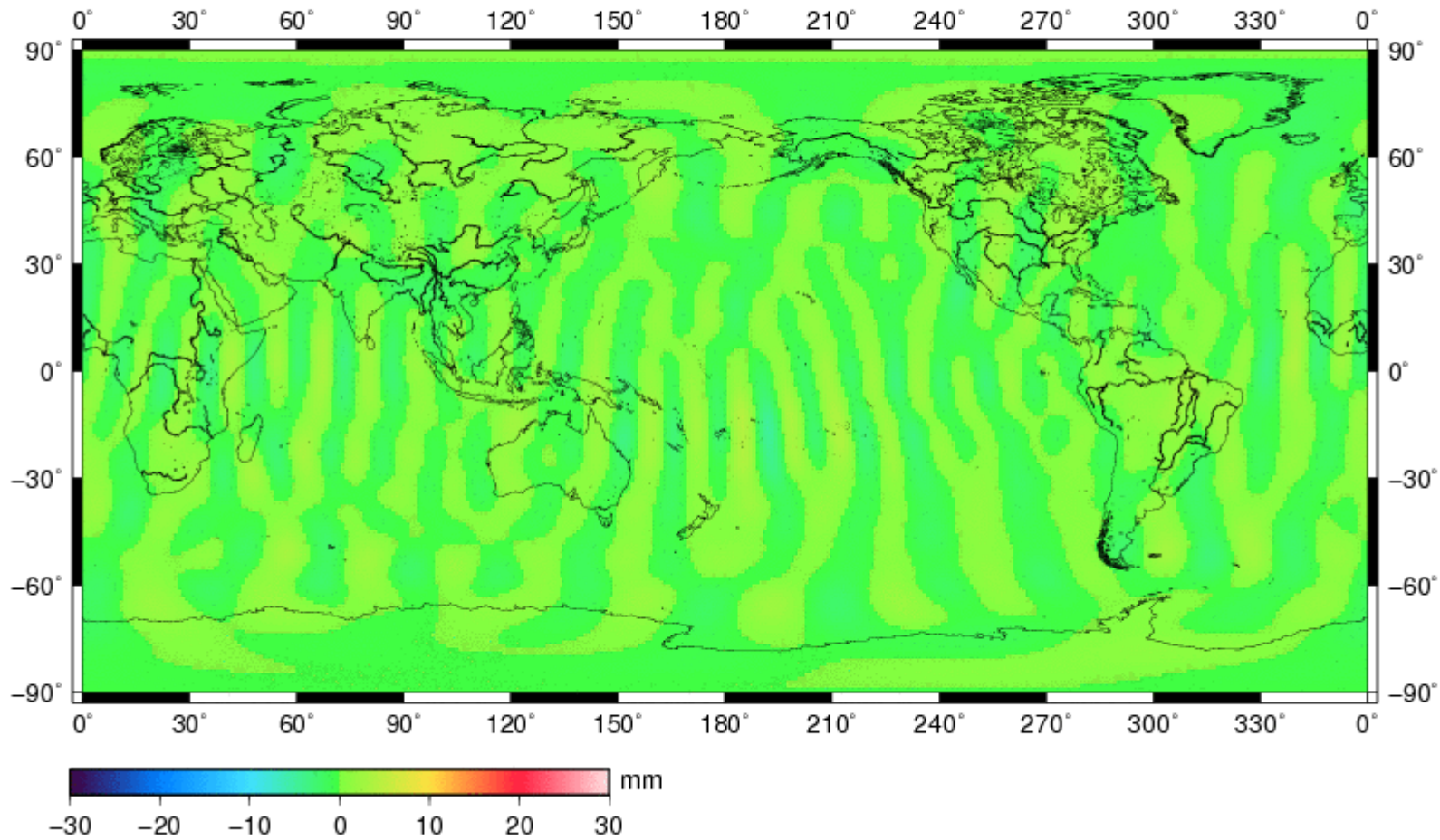
## Sum of PC 2-10, spectral MSSA L=10

2002.121.139 PC\_sum\_1

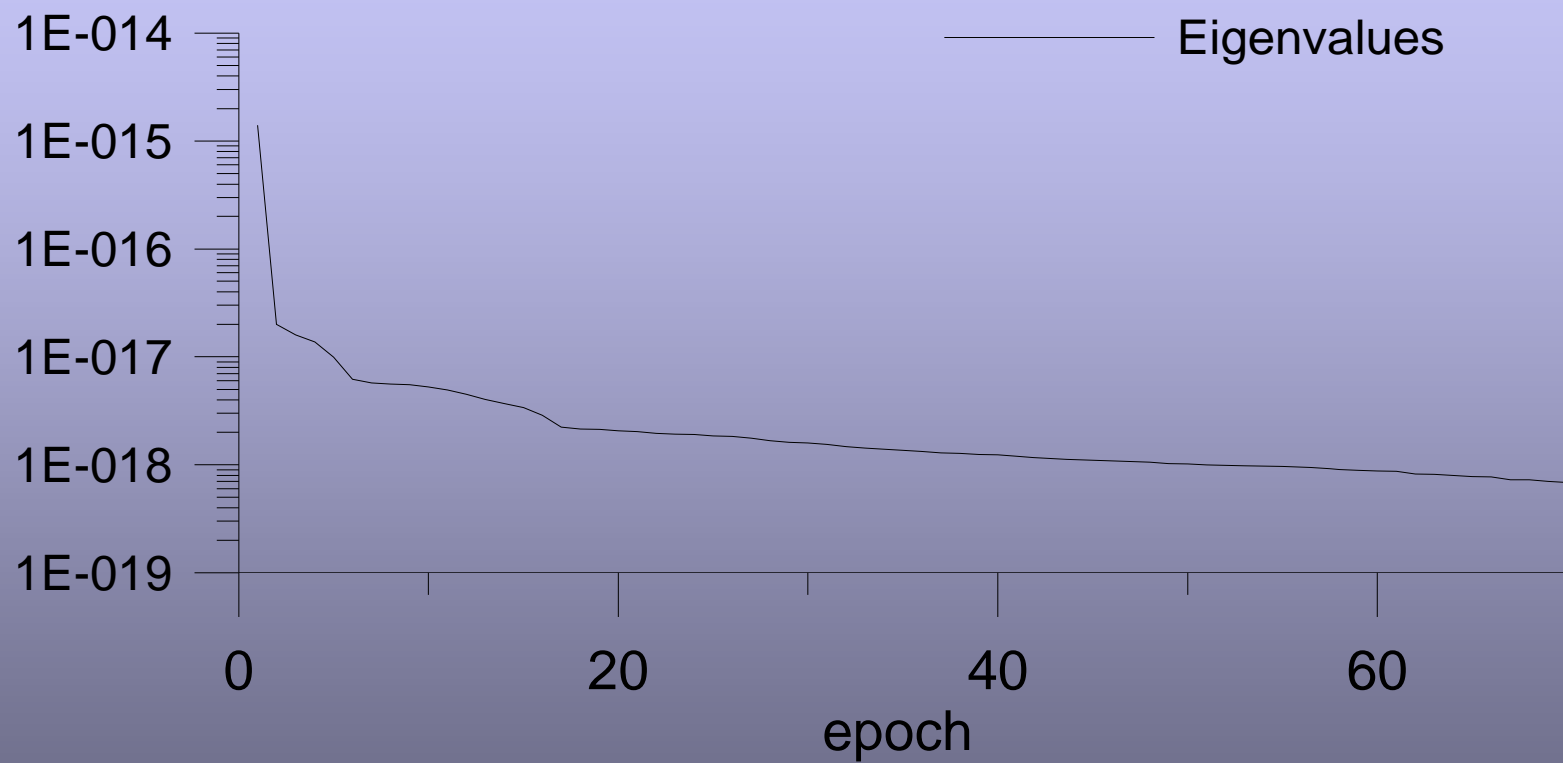


# Difference between initial signal and sum of PC 1-10, spectral MSSA L=10

2002.121.139 PC\_diff\_1

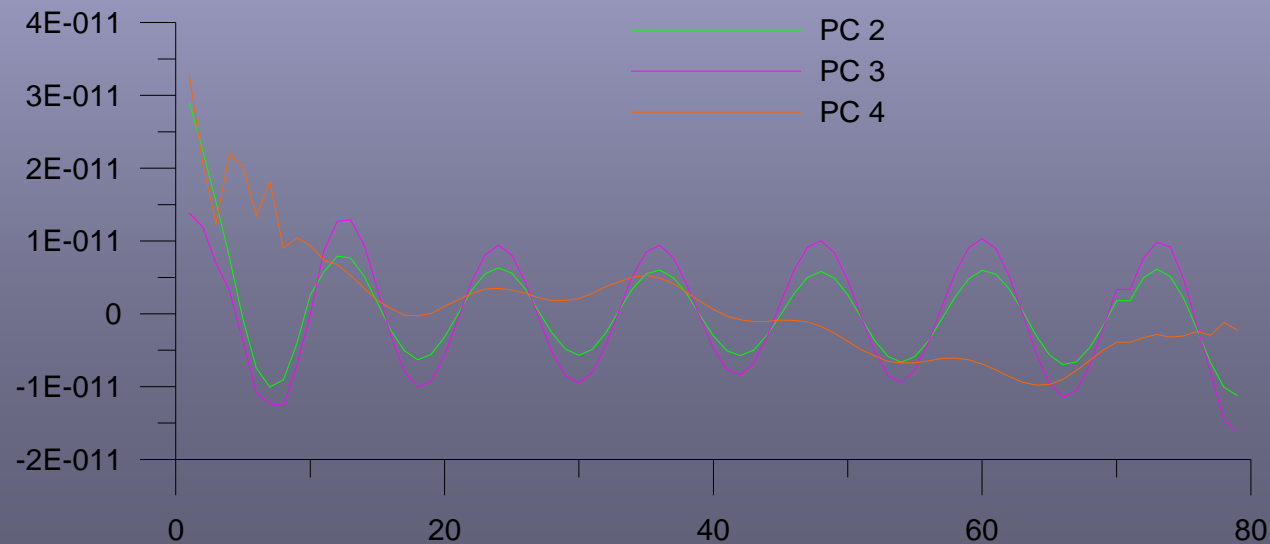
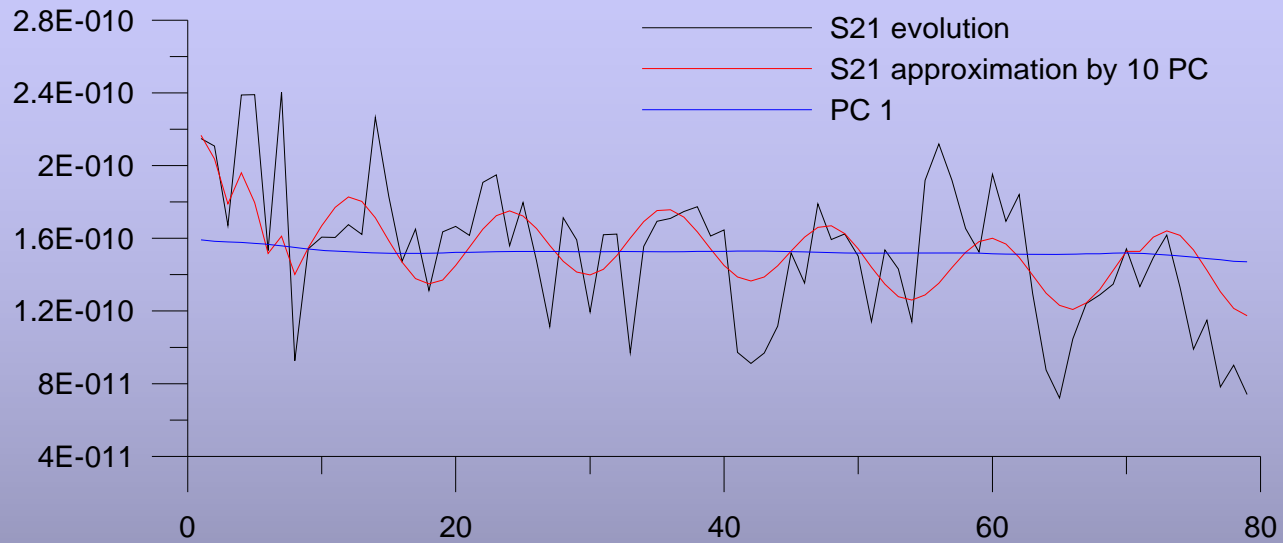


# EIGENVALUES





# Approximation of Stokes coefficient $S_{21}$ by principal components of MSSA, L=10



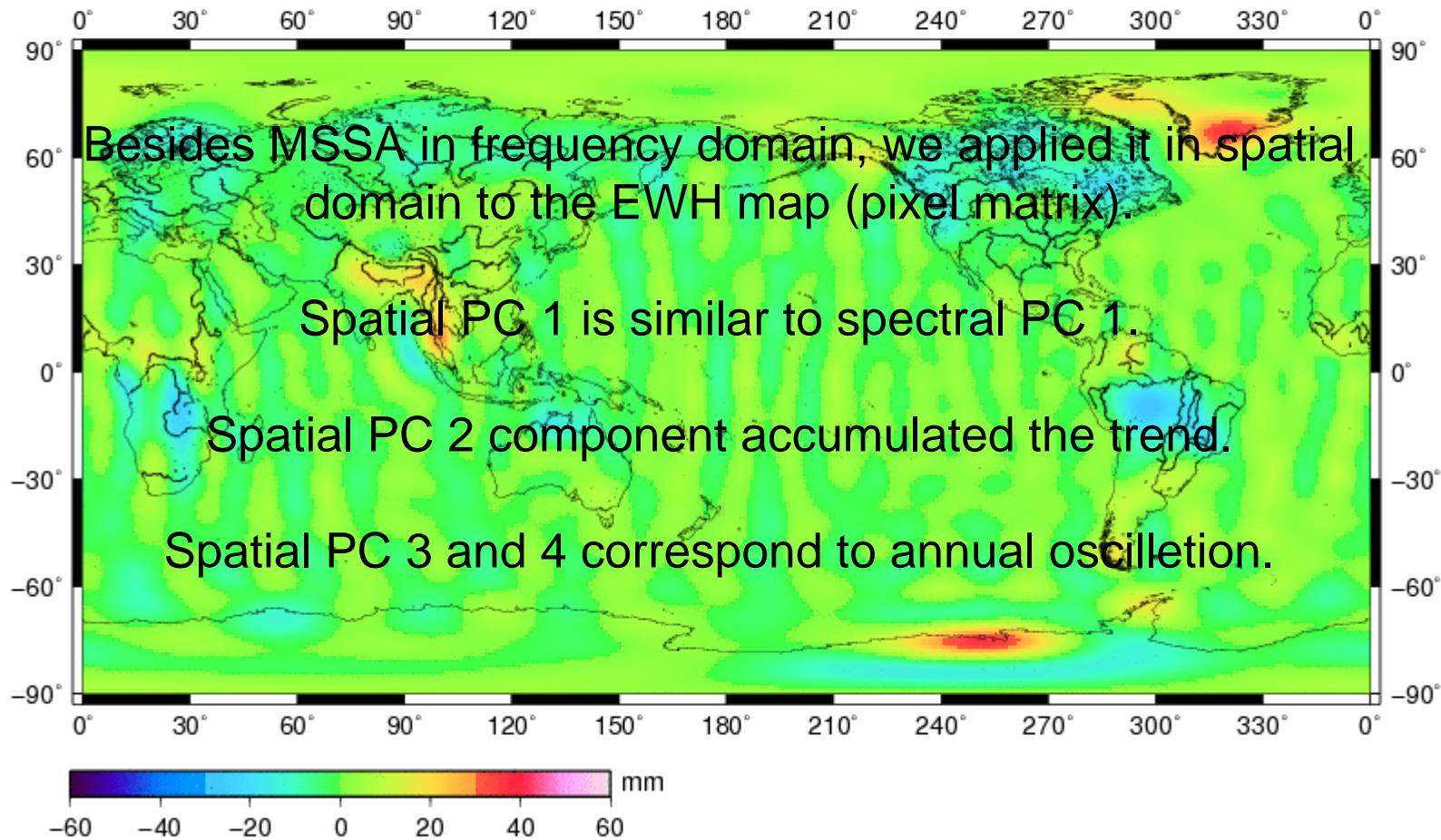
2002

2006

2009

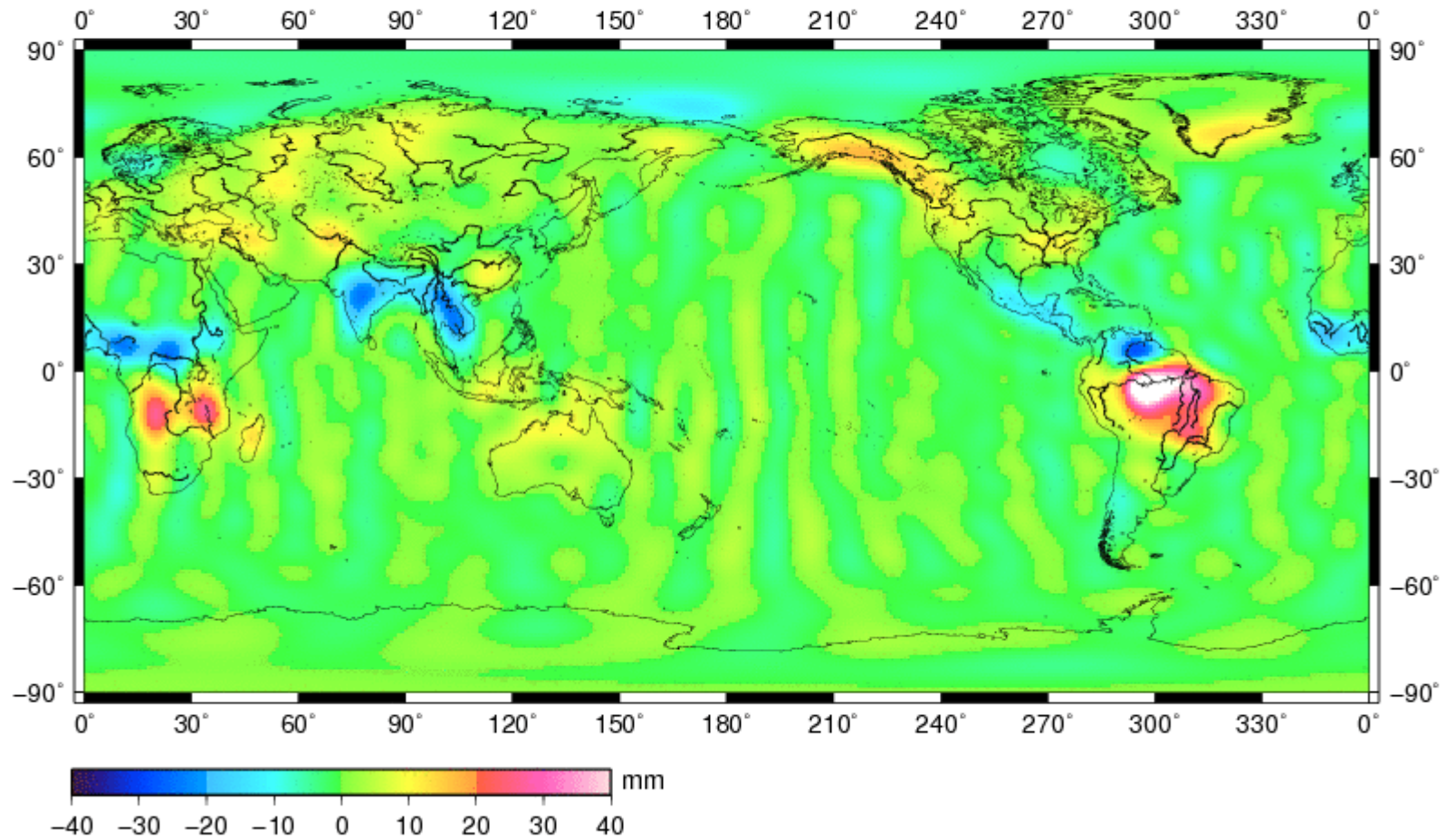
## PC 2 for MSSA in spatial domain, L=10

2002.121.139 PC\_2\_1



## PC 3,4 for spatial MSSA, L=10

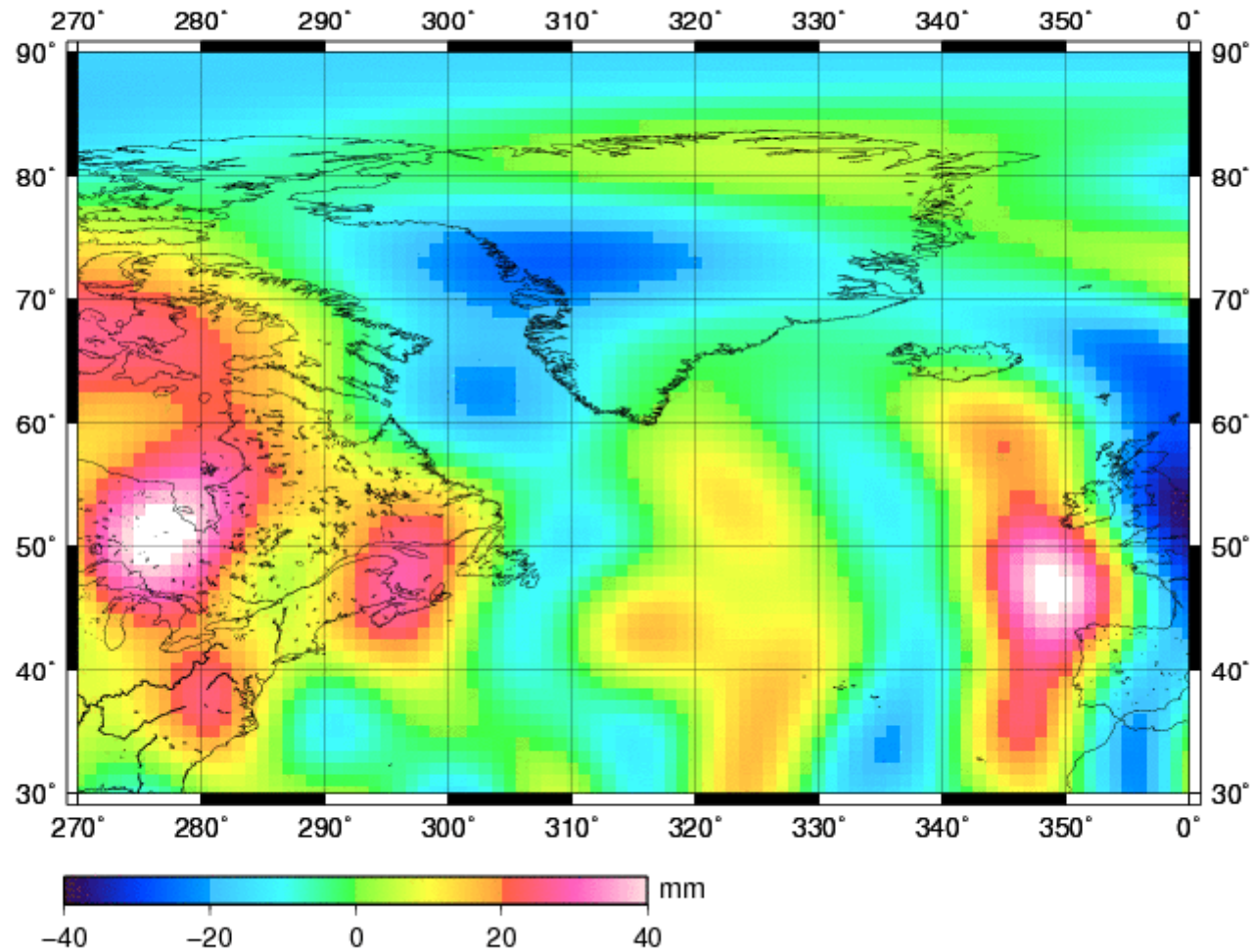
2002.121.139 PC\_4\_1





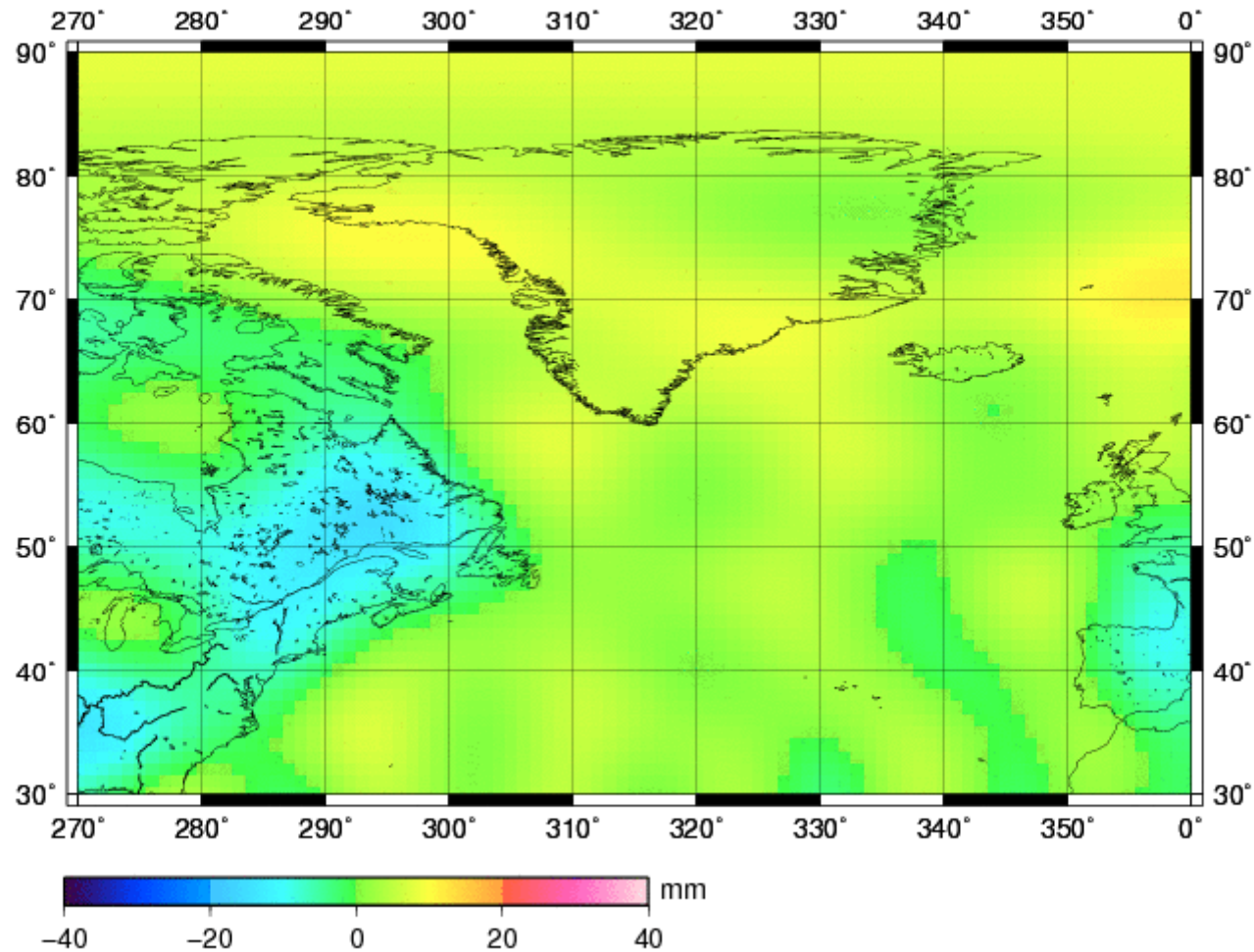
# Greenland PC 1 spatial MSSA, L=10

2002.121.139 PC\_1\_1



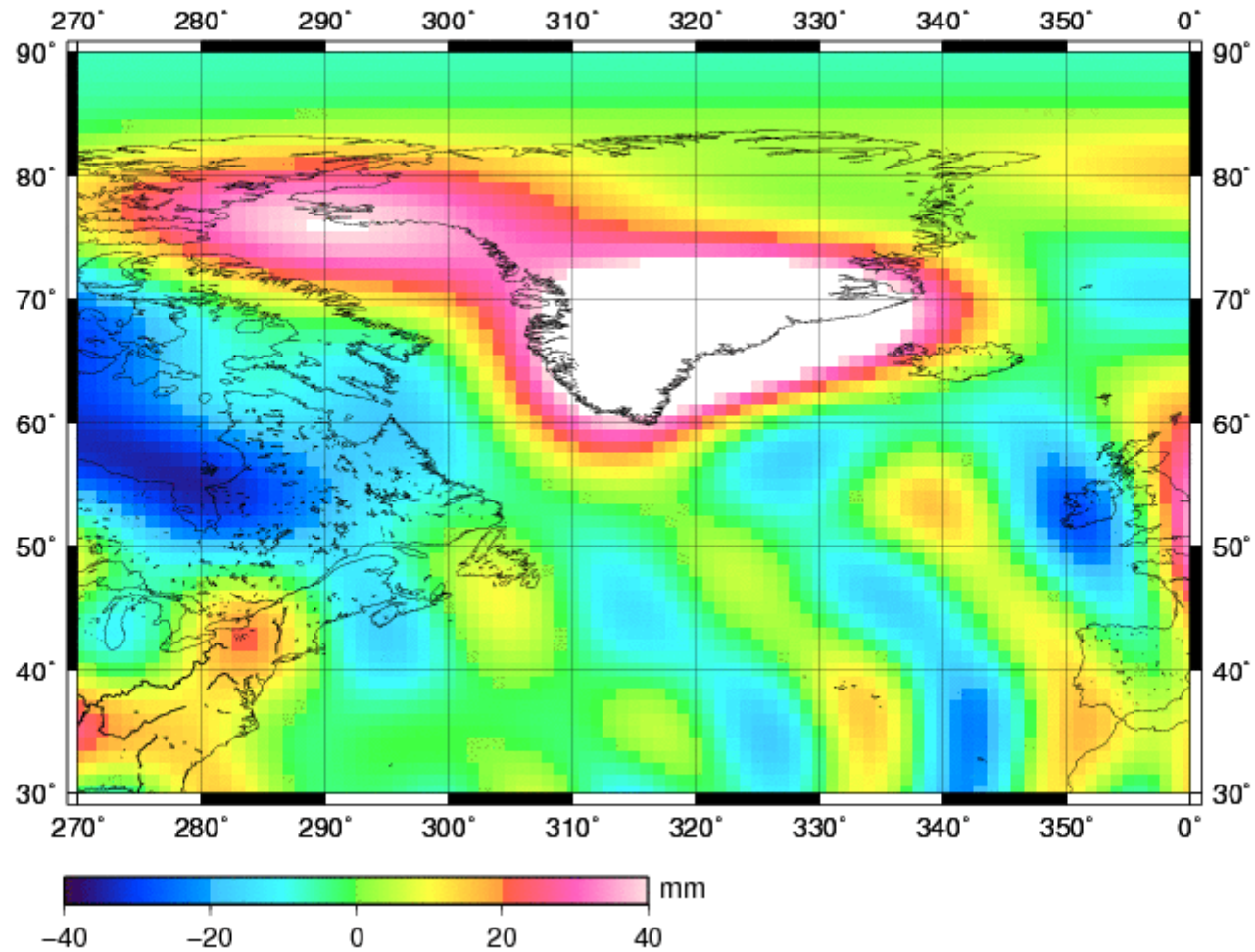
# Greenland PC 2 spatial MSSA, L=10

2002.121.139 PC\_2\_1



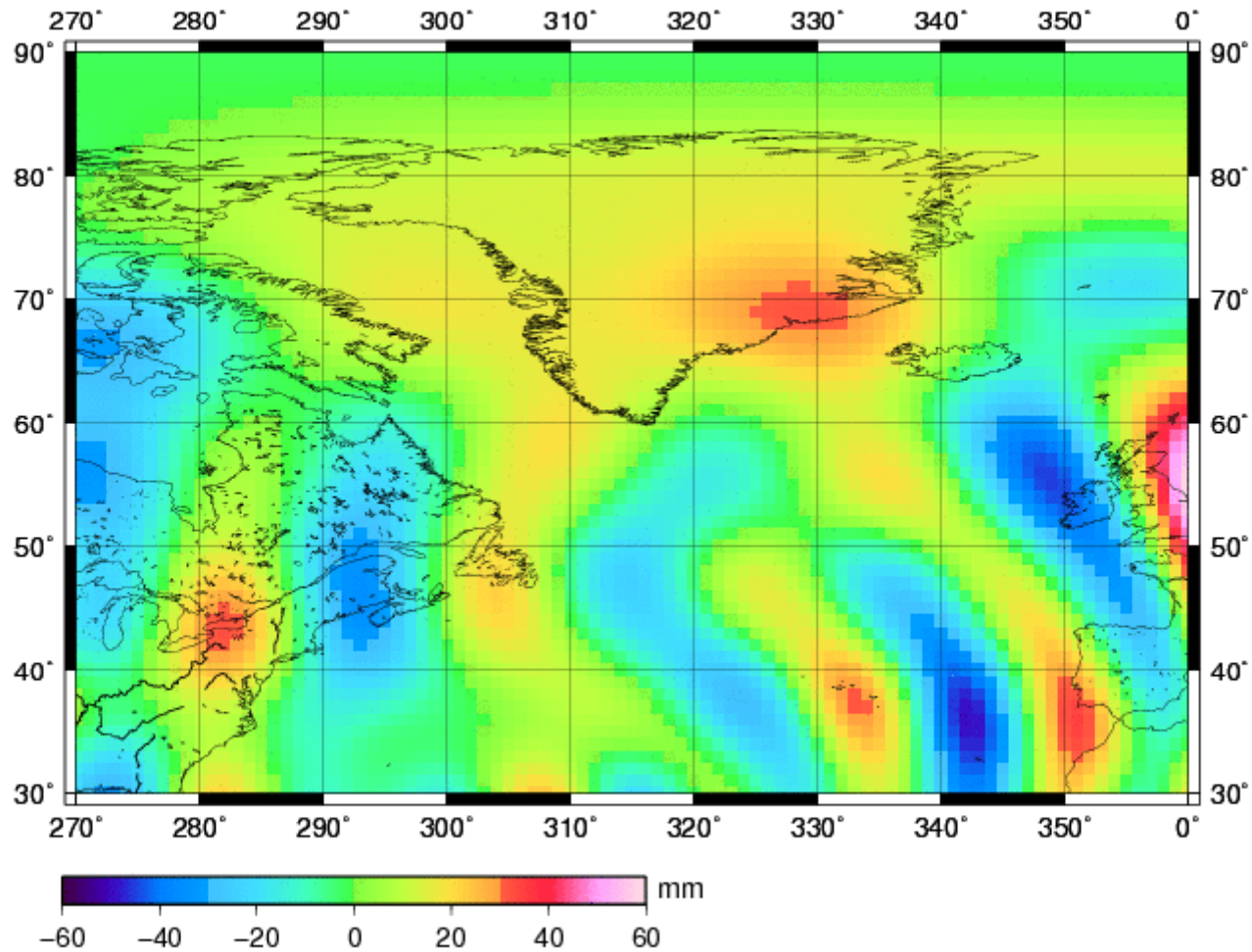
# Greenland PC 3,4 spatial MSSA, L=10

2002.121.139 PC\_4\_1



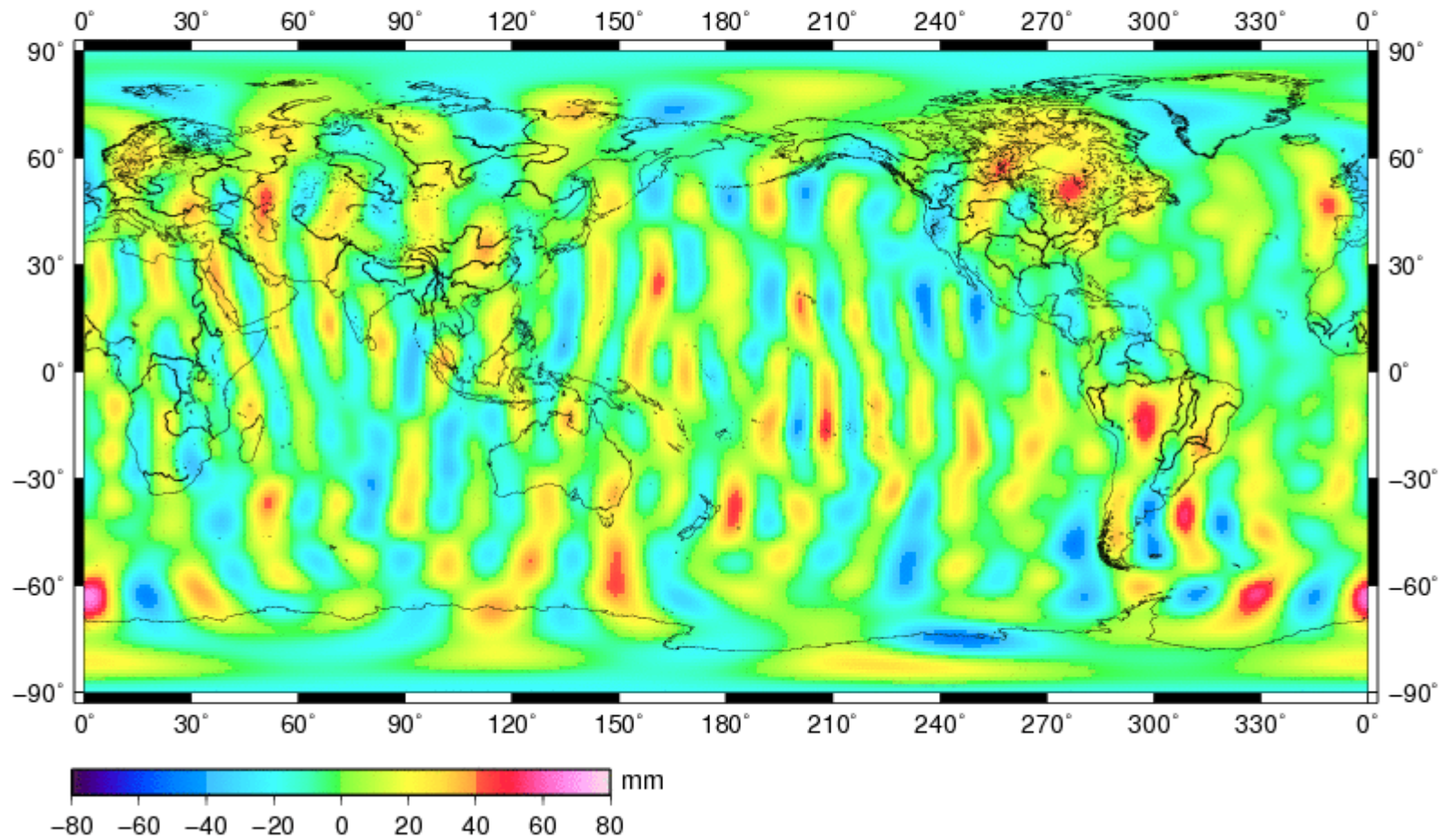
# Greenland PC 2-10 spatial L=10

2002.121.139 sum 2-10\_1



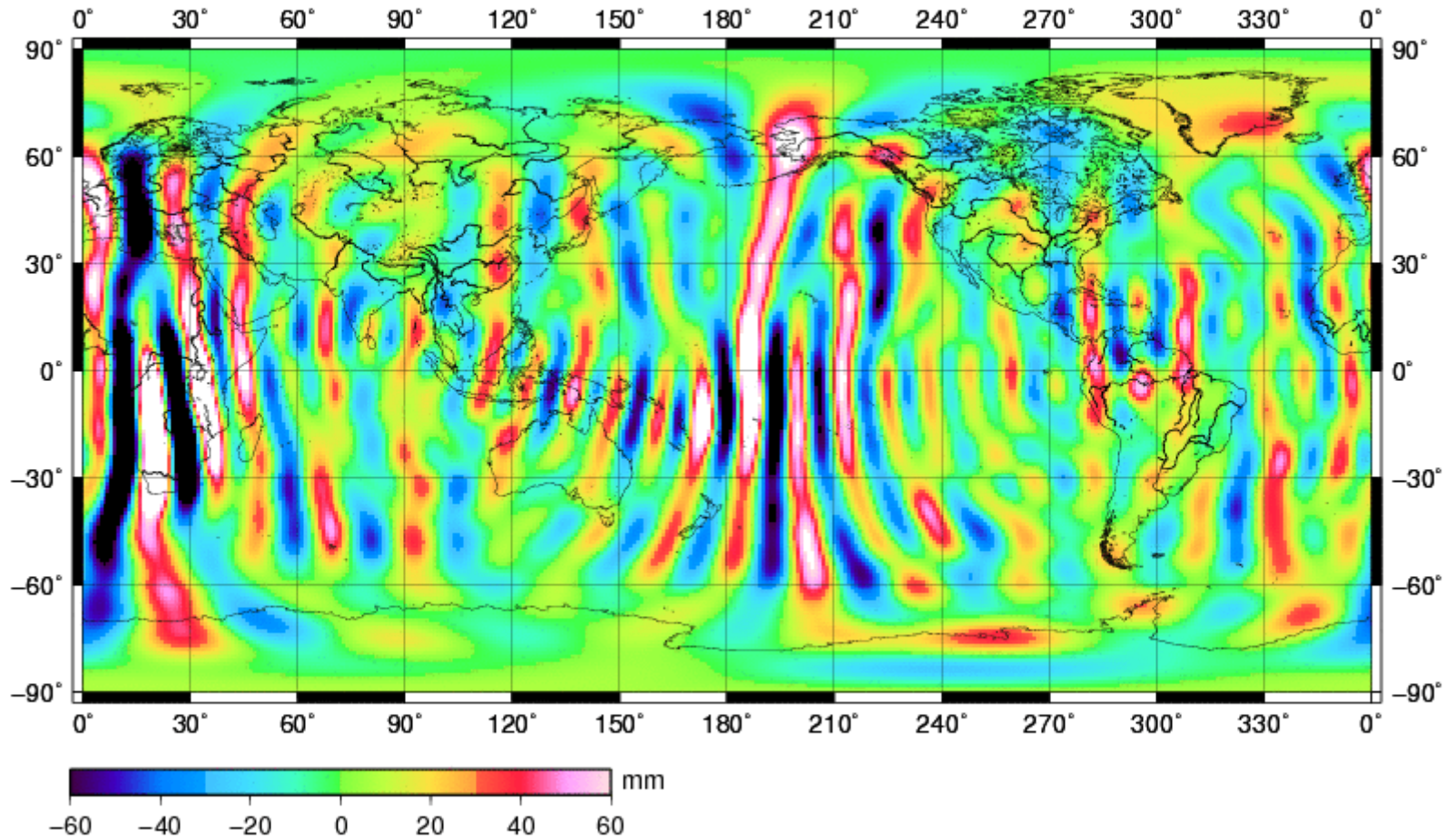
# PC 1 spatial PCA, L=1

2002.121.139 PC\_1\_1



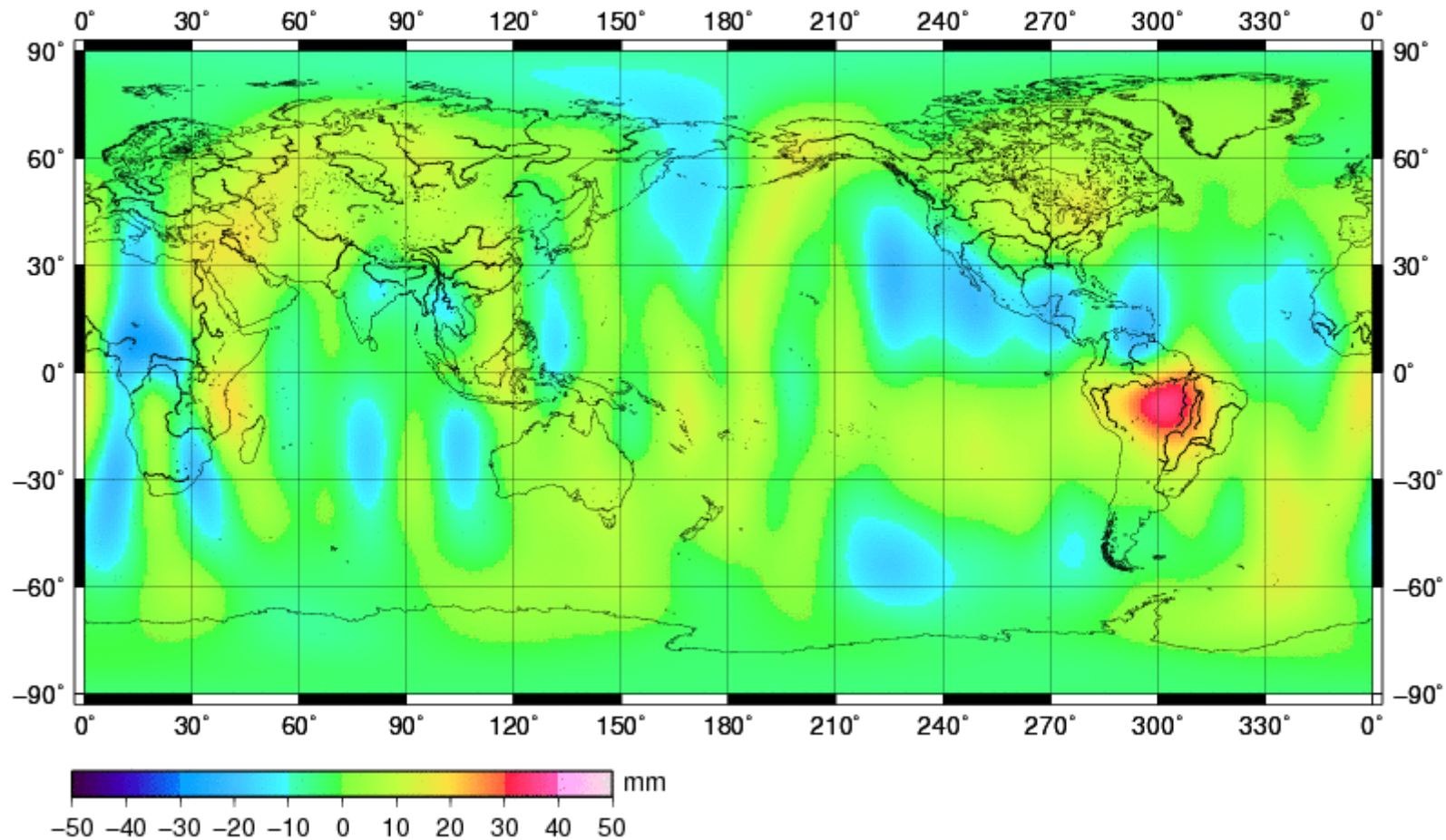
# PC 2,3 spectral PCA, L=1

2002.121.139 PC 2+3\_1



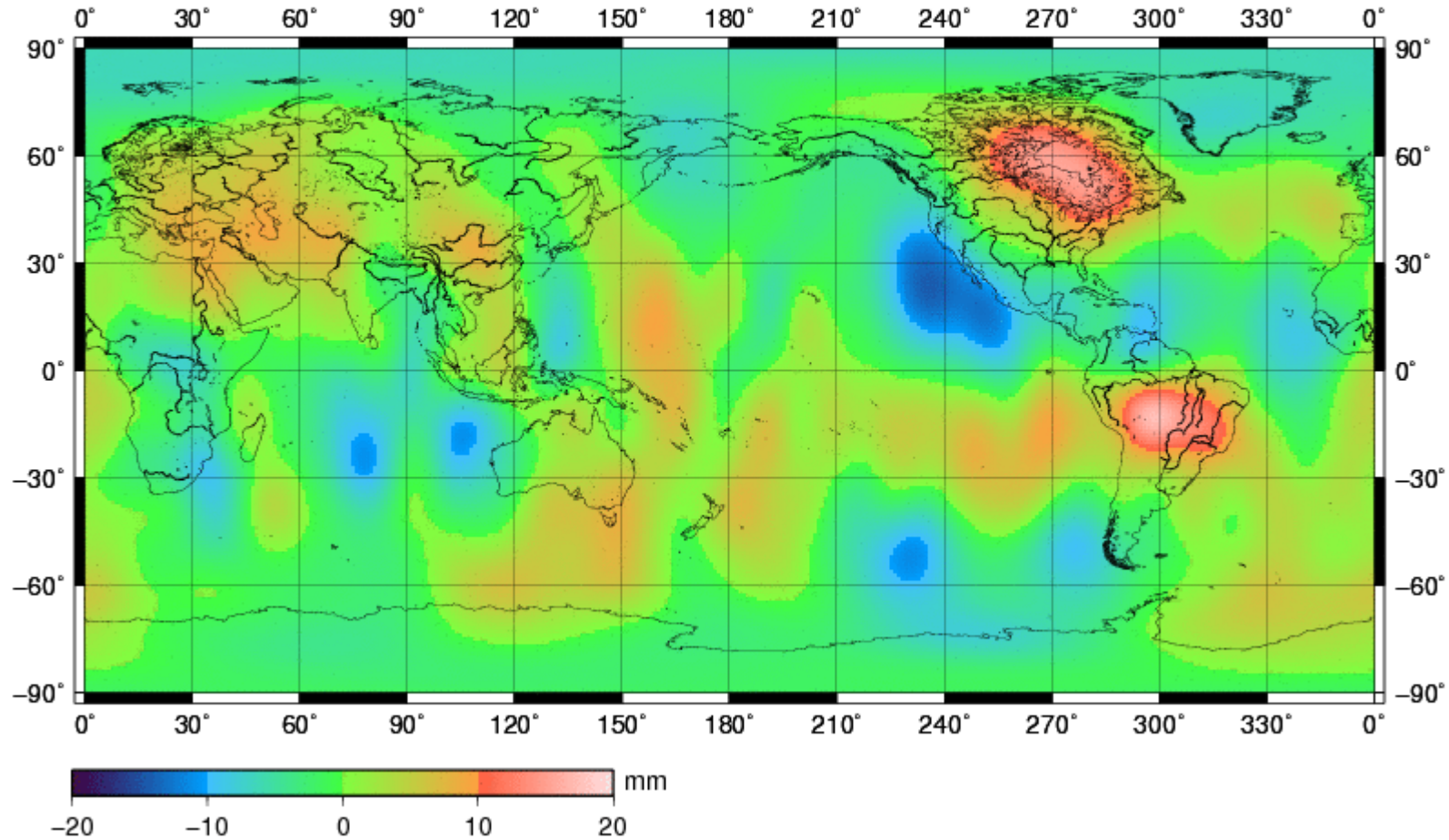
# INITIAL FILTERED DATA (Gaussian filter R=800km)

2002.121.139 init\_2



# PC 1 spatial filtered L=10

2002.121.139 PC\_1\_2





## Panteleev corrective smoothing

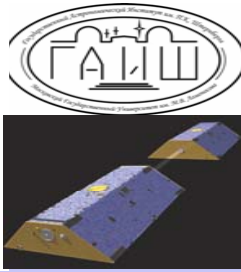
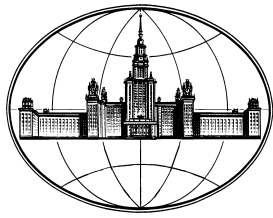
If we have signal  $u$ , which passes through the dynamic system (equipment) with transfer function  $W$  and gives us observations  $f$  together with noise, we can reconstruct  $u$  applying inverse operator together with smoothing operator  $W_{smooth}$ . Smoothing is very important to prevent amplification of noises. If dynamic system is linear, we can write down the corrective smoothing in spectral domain (hat means Fourier transform) as

$$\hat{u}_{corrected} = \frac{W_{smooth}(i\omega)}{W(i\omega)} \hat{f}(\omega) = W_{Pantelev}(i\omega) \hat{f}(\omega)$$

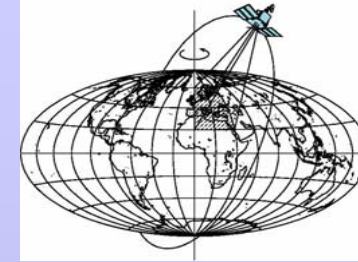
Method can be used instead of regularization. We can extract useful information from the signal only if we are able to find basis (functional, vectorial, etc.), where it is separable from the undesirable systematic errors and noise. If we can use SSA basis for denoising and realize  $W_{smooth}$  in this basis, we would be able to reconstruct the signal  $u$  successfully. At that, we can apply  $W_{smooth}$  before or after inversion. If we have been able to decompose Level-2 signal, as it's shown, we should be able to decompose also the initial unprocessed signal and use this approach for inversion.

## References

1. Rangelova E. et al., Analysis of GRACE time-variable mass redistribution signals over North America by means of principal components analysis, JGR, Vol. 112, 2007;
2. Wouters B., Schrama E., Improved accuracy of GRACE gravity solution through empirical orthogonal function filtering of spherical harmonics. GRL Vol. 34, 2007
3. Golyandina N. et al., Analysis of time series structure: SSA and related techniques, Chapman & Hall, 2001
4. Jolliffe I.T., Principal Component Analysis, Springer, 2001
5. Panteleev V.L., Algorithms of smoothing of aerogravimetric observations, correcting dynamic errors of the measurements. Izvestiya RAS, Physics of the solid Earth, N3, 2001 (in Russian)



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