

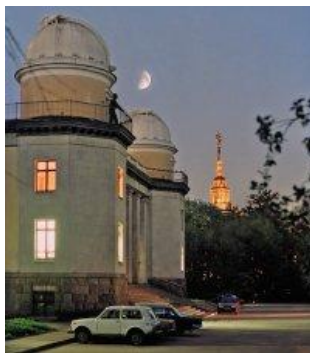
Regularized Kalman Filter in application to the Chandler wobble of the pole

Leonid Zotov^{1,6},

C. Bizouard², N. Sidorenkov³, C.K. Shum⁴, S. Denisenko⁵
wolftempus@gmail.com

¹SAI MSU, Russia ²Paris Observatory ³Hydrometcenter of Russia

⁴School of Earth Sciences OSU ⁵Snt Petersburg University ⁶MIEM HSE



ICCC workshop
“Geodesy for climate research”
29 March 2023



Kalman filter as least squares

Let's start from observation equation

$$y = Cx + r$$

where y is N -dimensional, x is n -dimensional vectors and recurrent state equation

$$x_{k+1} = Ax_k + Bu_k$$

with $A - [n \times n]$, $B - [n \times m]$ matrices, u_k is m -dimensional input white noise. We would require

$$\|y - Cx\| \rightarrow \min \quad (5)$$

and

$$\|\hat{x}_{k+1} - A\hat{x}_k\| \rightarrow \min \quad (6)$$

Kalman filter as least squares

allows to write the matrix system for this optimisation problems in the form of least squares

$$\begin{bmatrix} A\hat{x}_k \\ y_{k+1} \end{bmatrix} = \begin{bmatrix} I \\ C_{k+1} \end{bmatrix} x_{k+1} + \begin{bmatrix} -A\epsilon_k - Bu_k \\ r_{k+1} \end{bmatrix}$$

Let's rewrite above equation using new h, H, s notations for vectors and matrices

$$h = Hx + s$$

Least squares solution is

$$\hat{x}_{k+1} = (H^T S H)^{-1} H^T S h$$

with weight matrix

$$S = \begin{Bmatrix} D_{k+1}^{*-1} & 0 \\ 0 & Q^{-1} \end{Bmatrix}$$

where inverted covariance matrices were used

$$D_{k+1}^* = AD_\epsilon A^T + BD_u B^T$$

$$Q = D_r = \sigma_0^2 I$$

$$H^T S H = D_{k+1}^{-1} = D^{*-1} + \sigma_0^{-2} C^T C$$

$$H^T S h = D_{k+1}^{-1} \hat{x}_{k+1}^* + \sigma_0^{-2} C^T y_{k+1}$$

then

$$\hat{x}_{k+1} = (D^{*-1} + \sigma_0^{-2} C^T C)^{-1} (D^{*-1} \hat{x}_{k+1}^* + \sigma_0^{-2} C^T y_{k+1}) \quad (7)$$

Regularization

Imagine that matrix C of observational equation is rank-deficient or close to this, then the normal matrix will be ill-posed and LS solution ill-conditioned.

Tykhonov regularization approach recommends to search for normal pseudo-solution by minimizing

$$L_{Tykhonov} = \|y - Cx\| + \alpha \|x\| \rightarrow \min \quad (11)$$

For least squares

$$\hat{x} = (C^T PC + \alpha I)^{-1} C^T P y$$

$$D_{\hat{x}} = (C^T PC + \alpha I)^{-1}.$$

For recurrent least squares

$$\hat{x}_{N+1} = (C^T PC + a^T p a + \alpha I)^{-1} (C^T P y + a^T p y_{N+1}) \quad (13)$$

denoting

$$D_N^{-1} = (C^T PC + \alpha I), \quad D_{N+1}^{-1} = (C^T PC + a^T p a + \alpha I)$$

Sequential estimate does not change

$$\hat{x}_{N+1} = \hat{x}_N + K p^{-1} (y_N - a \hat{x}_N).$$

For covariance matrix update we obtain (see above)

$$D_{N+1} = D_N - K \frac{a}{\sqrt{p}} D_N$$

Kalman filter regularization

$$\hat{x}_{k+1} = (H^T S H + \alpha I)^{-1} H^T S h = \underbrace{(D^{*-1} + C^T Q^{-1} C + \alpha I)^{-1}}_{(17)} (D^{*-1} \hat{x}_{k+1}^* + C^T Q^{-1} y_{k+1})$$

Recursive expansion of

$$K = D_{k+1}^{reg} C^T Q^{-1}$$

gives
$$K = (Q + C D^* C^T)^{-1} (D^* C^T - \alpha D D^* C^T). \quad 18$$

which only under some simplifications can be brought to the widely-used approach 19

Observational error covariance matrix Q can play a role of regularization parameter.

So, we finally use this approach.

$$\begin{aligned} x_k &= x_k^* + K_k (y_k - C_k x_k^*), \\ K_k &= D_k^* C_k^T [C_k D_k^* C_k^T + Q_k + \alpha I]^{-1}, \quad 19 \\ D_k &= D_k^* - K_k C_k D_k^*, \\ x_{k+1}^* &= A_k x_k. \end{aligned}$$

3.1 Tikhonov Regularized Kalman Filter and Its Algorithm

The objective function of TRKF is designed as follows

$$\Psi(\hat{X}_k^{TRKF}) = V_{\hat{x}_k^{TRKF}}^T R_k^{-1} V_{\hat{x}_k^{TRKF}} + V_{\hat{x}_k^{TRKF}}^T P_{k/k-1}^{-1} V_{\hat{x}_k^{TRKF}} + \alpha_k Z_k^T H_k Z_k \quad (14)$$

where $\alpha_k > 0$ is the regularization parameter, $H_k > 0$ is the regularization matrix and $Z_k = \hat{X}_k^{TRKF} - \hat{X}_{k/k-1}$.

Using the weighted least square method, we can get

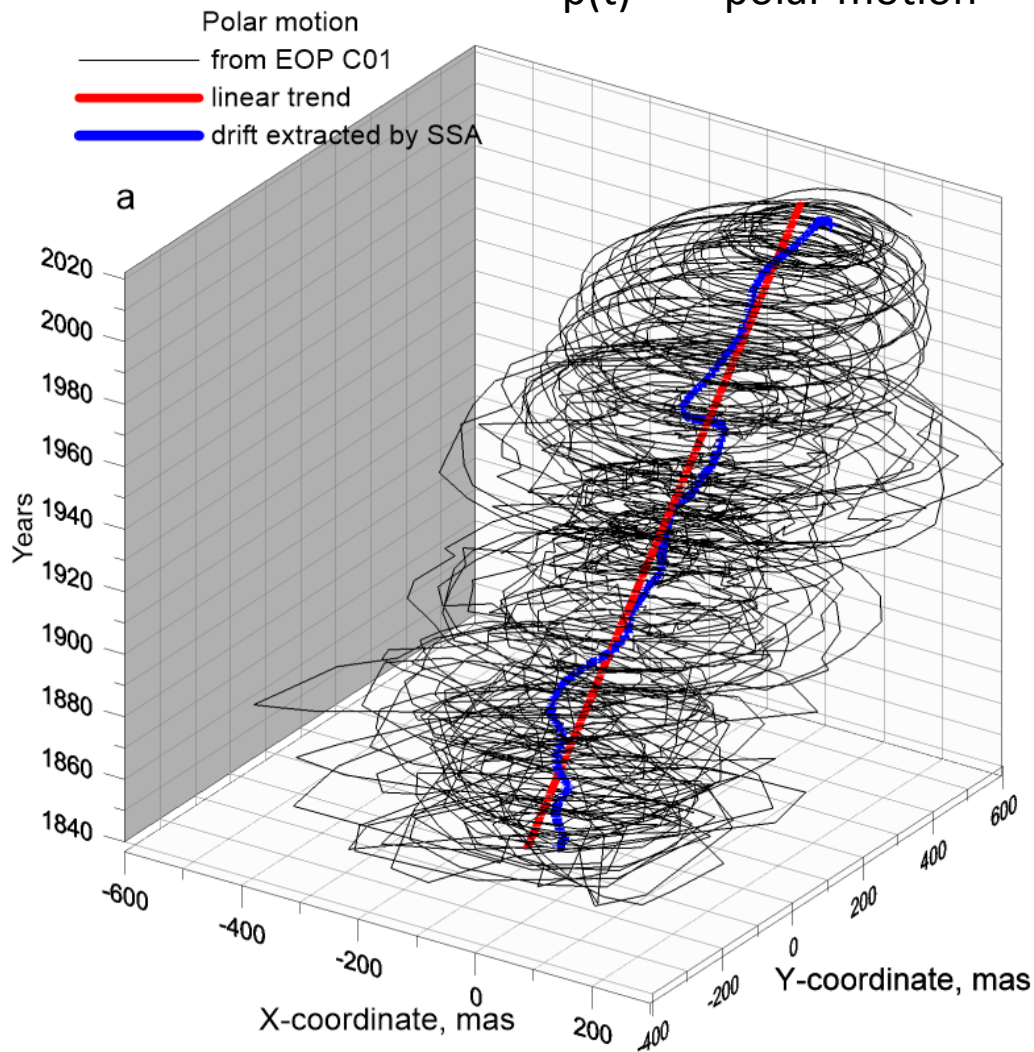
$$K_k^{TRKF} = (A_k^T R_k^{-1} A_k + P_{k/k-1}^{-1} + \alpha_k H_k)^{-1} A_k^T R_k^{-1} \quad (15)$$

$$\hat{X}_k^{TRKF} = \hat{X}_{k/k-1} + K_k^{TRKF} (L_k - A_k \hat{X}_{k/k-1}) \quad (16)$$

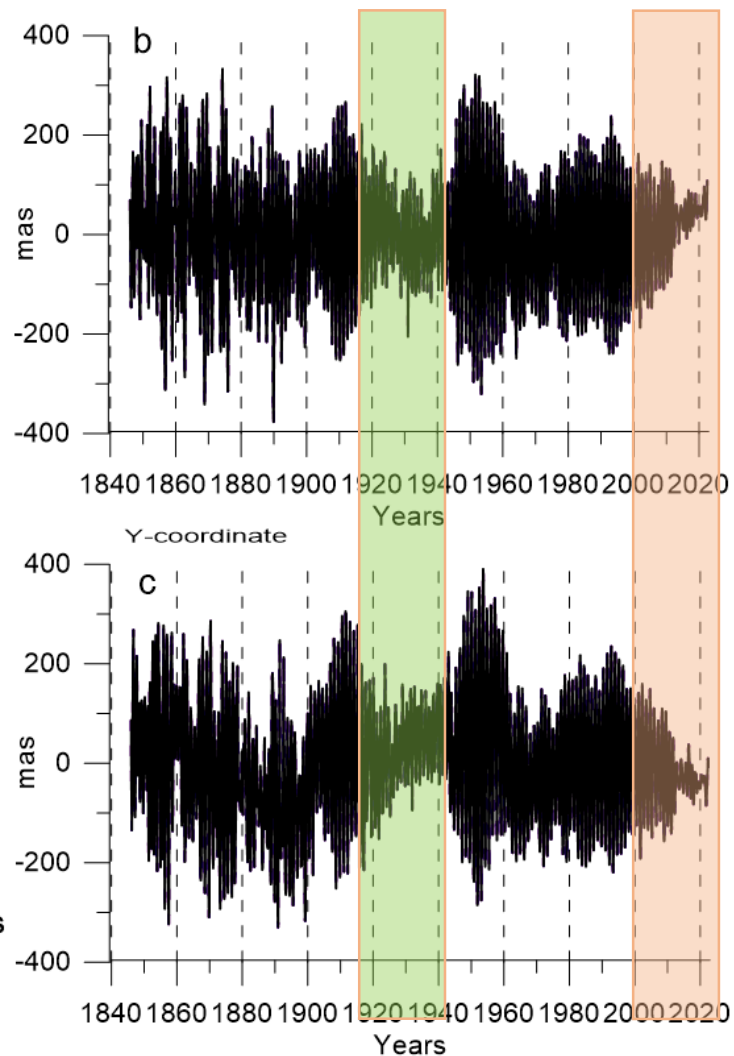
$$P_k^{TRKF} = (I - K_k^{TRKF} A_k) P_{k/k-1} \quad (17)$$

Polar motion since 1846 from EOP C01 bulletin

$p(t)$ - polar motion



X-coordinate of the pole after trend and annual mode were subtracted
— from EOP C01 since 1846 yr



DYNAMICAL MODEL OF THE POLAR MOTION

$$\frac{i}{\sigma_c} \frac{dp(t)}{dt} + p(t) = \chi(t)$$

$$p = p_1 + ip_2$$

$$\chi = \chi_{mass} + \chi_{motion}$$

$$\sigma_c = 2\pi f_c (1 + i/2Q)$$

$$f_c = \frac{1}{433} \text{ days}^{-1} \quad Q = 100$$

In time domain

$$\frac{-\omega \hat{p}(\omega)}{\sigma_c} + \hat{p}(\omega) = \hat{\chi}(\omega)$$

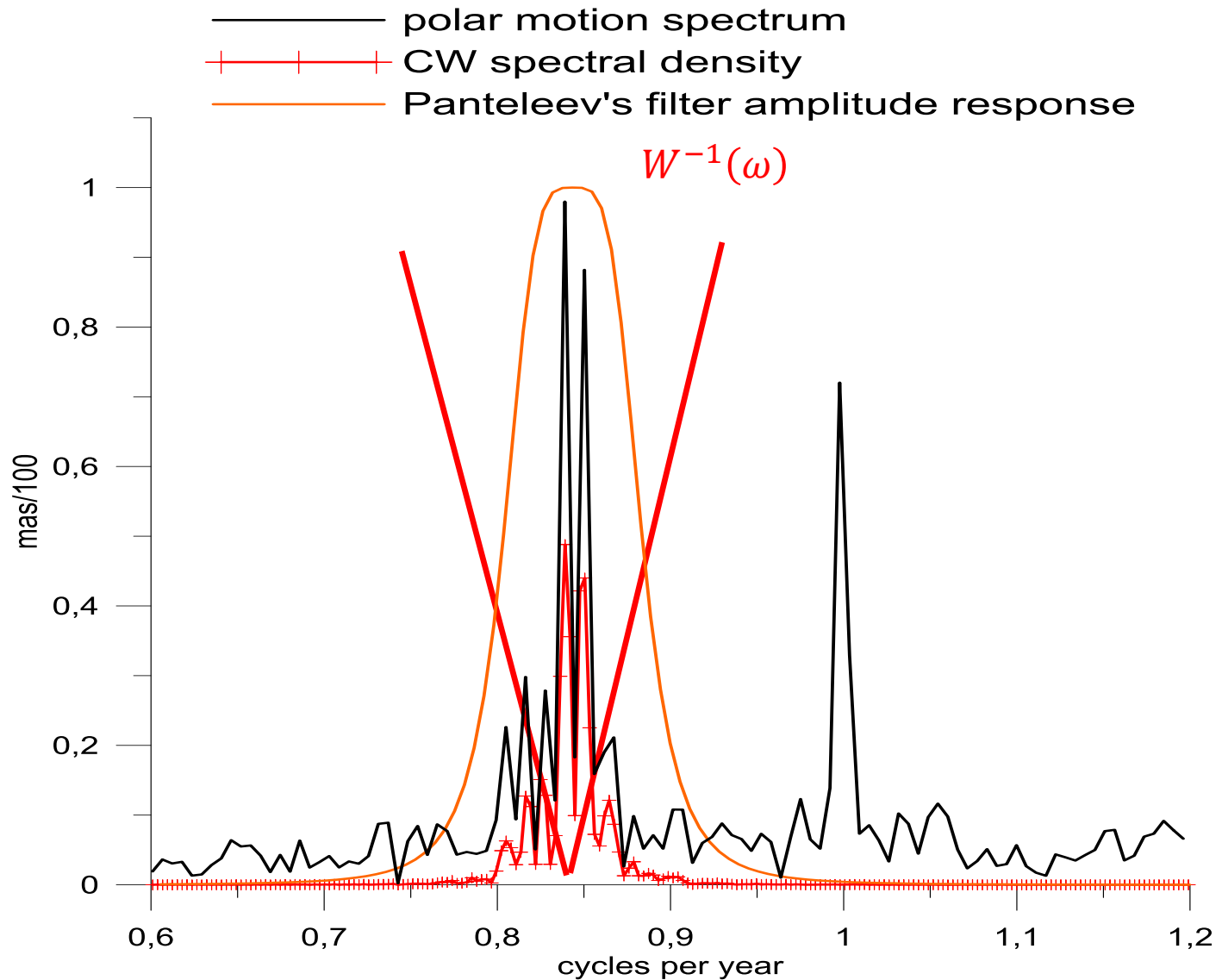
$$\frac{\sigma_c - \omega}{\sigma_c} \hat{p}(\omega) = \hat{\chi}(\omega)$$

$$\hat{p}(\omega) = W(\omega) \hat{\chi}(\omega)$$

$$W(\omega) = \frac{\sigma_c}{\sigma_c - \omega}$$

In frequency domain

Polar motion spectrum and Panteleev's corrective filtering



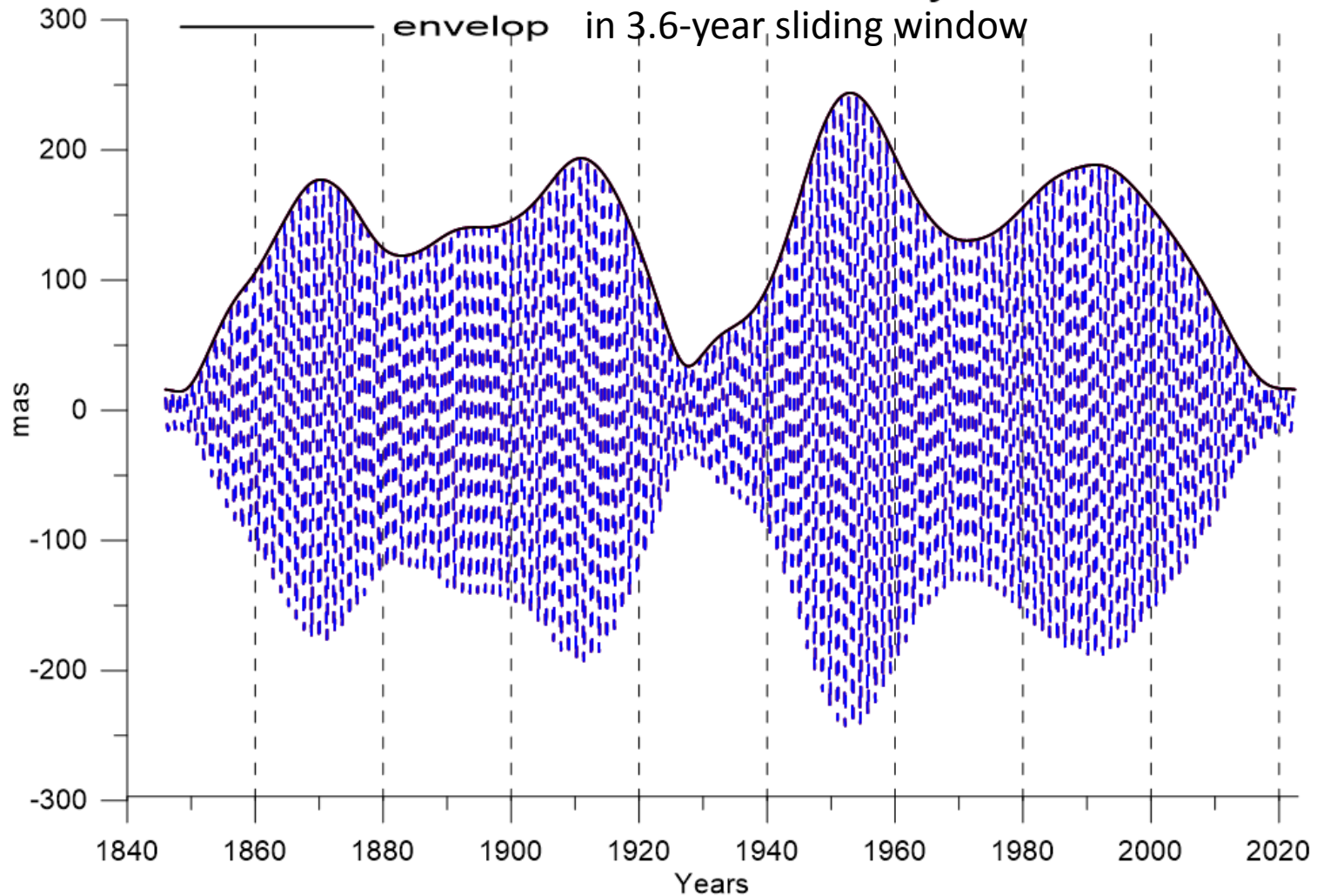
Changes of the Chandler wobble amplitude

$p_1(t)$ - polar motion

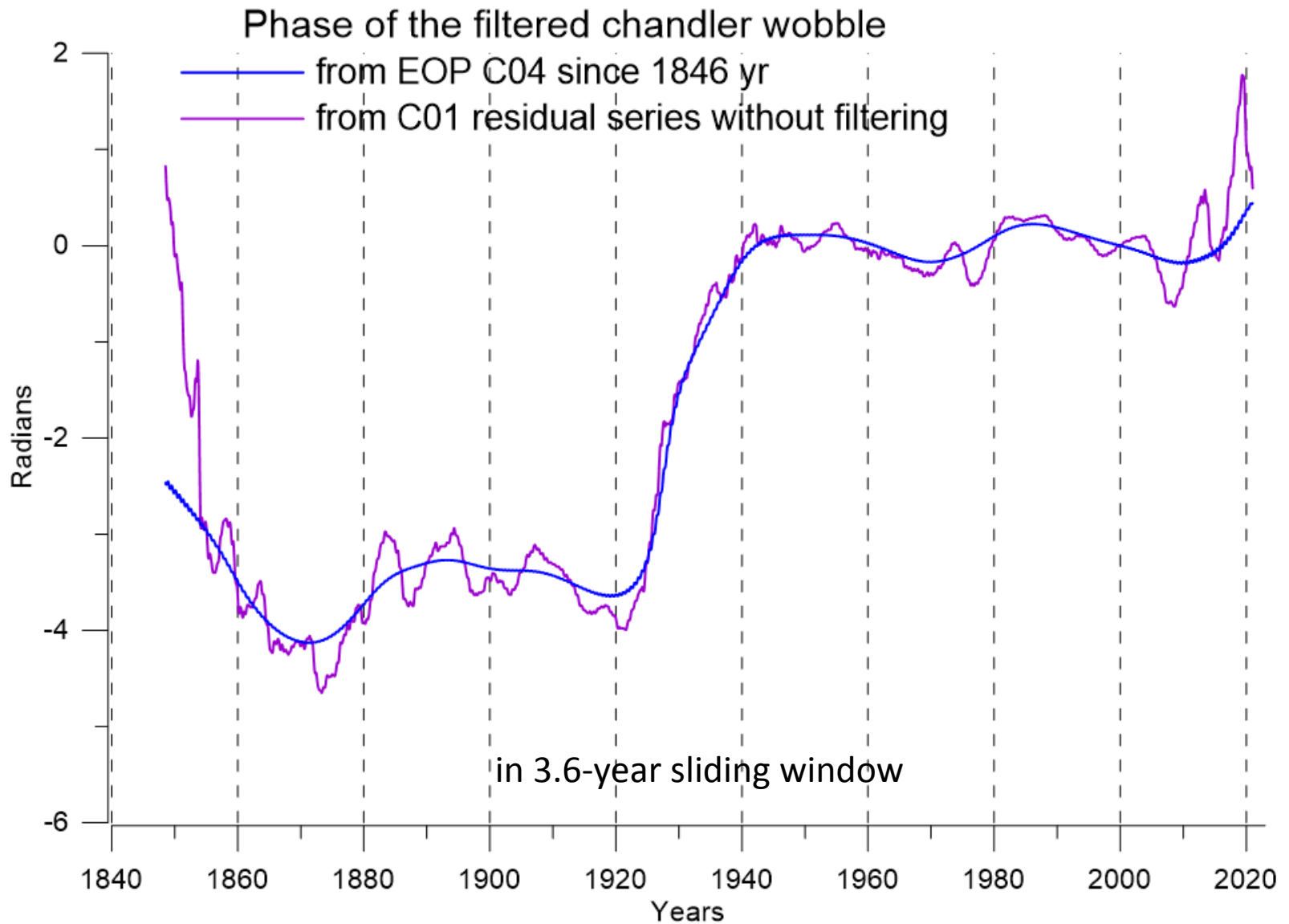
Filtered chandler wobble
(X-coordinate)

--- from EOP C01 since 1846 yr

— envelop in 3.6-year sliding window



Changes of the Chandler wobble phase



Matrix form of Euler-Liouville equations for the polar motion

State equation with coordinates of the pole $p_1 = x$ (along Greenwich) and $p_2 = -y$ (meridian 90° East) can be given in the form

$$D \begin{bmatrix} p_1 \\ p_2 \end{bmatrix} = \begin{bmatrix} -\beta & -\alpha \\ \alpha & -\beta \end{bmatrix} \begin{bmatrix} p_1 \\ p_2 \end{bmatrix} + \begin{bmatrix} \beta & \alpha \\ -\alpha & \beta \end{bmatrix} \begin{bmatrix} \chi_1 \\ \chi_2 \end{bmatrix},$$

Where α, β – are real and imaginary parts of complex-valued Chandler frequency $\sigma_c = \alpha + i\beta = 2\pi f_c + i\frac{\pi f_c}{Q}$, Q – quality factor (≈ 100), $f_c = 1/433 \text{ day}^{-1}$ – real Chandler frequency, $D =$

$D = \frac{d}{dt}$ – operator of differentiation,

χ_1, χ_2 – components of the input excitation, i – imaginary unit.

$$x = [p_1, p_2, \chi_1, \chi_2]^T \quad - ?$$

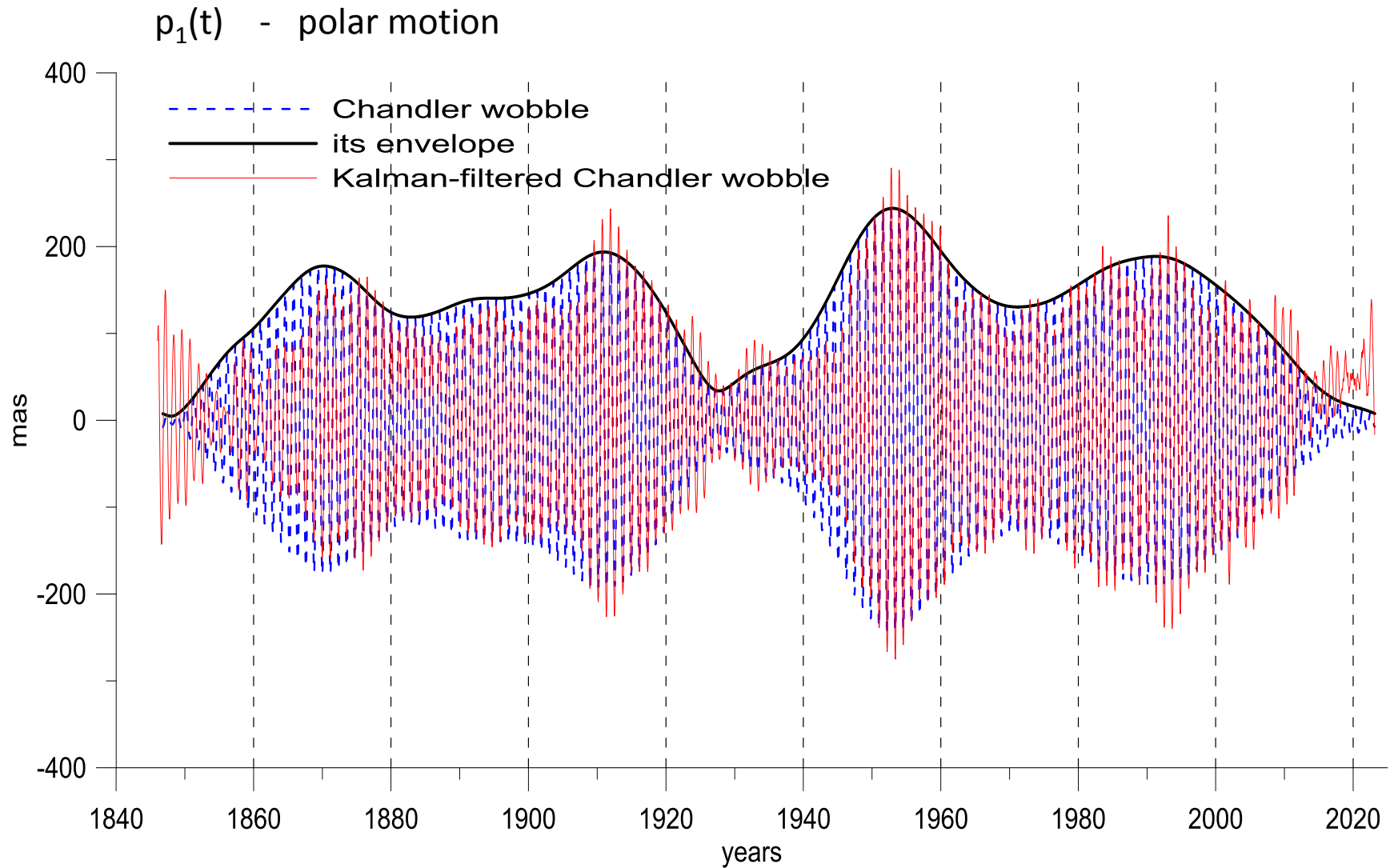
The problem of input excitation reconstruction is inverse

$$\begin{bmatrix} \chi_1 \\ \chi_2 \end{bmatrix} = \begin{bmatrix} \beta & \alpha \\ -\alpha & \beta \end{bmatrix}^{-1} \left(DI - \begin{bmatrix} -\beta & -\alpha \\ \alpha & -\beta \end{bmatrix} \right) \begin{bmatrix} p_1 \\ p_2 \end{bmatrix} = W^{-1}(D) \begin{bmatrix} p_1 \\ p_2 \end{bmatrix}$$

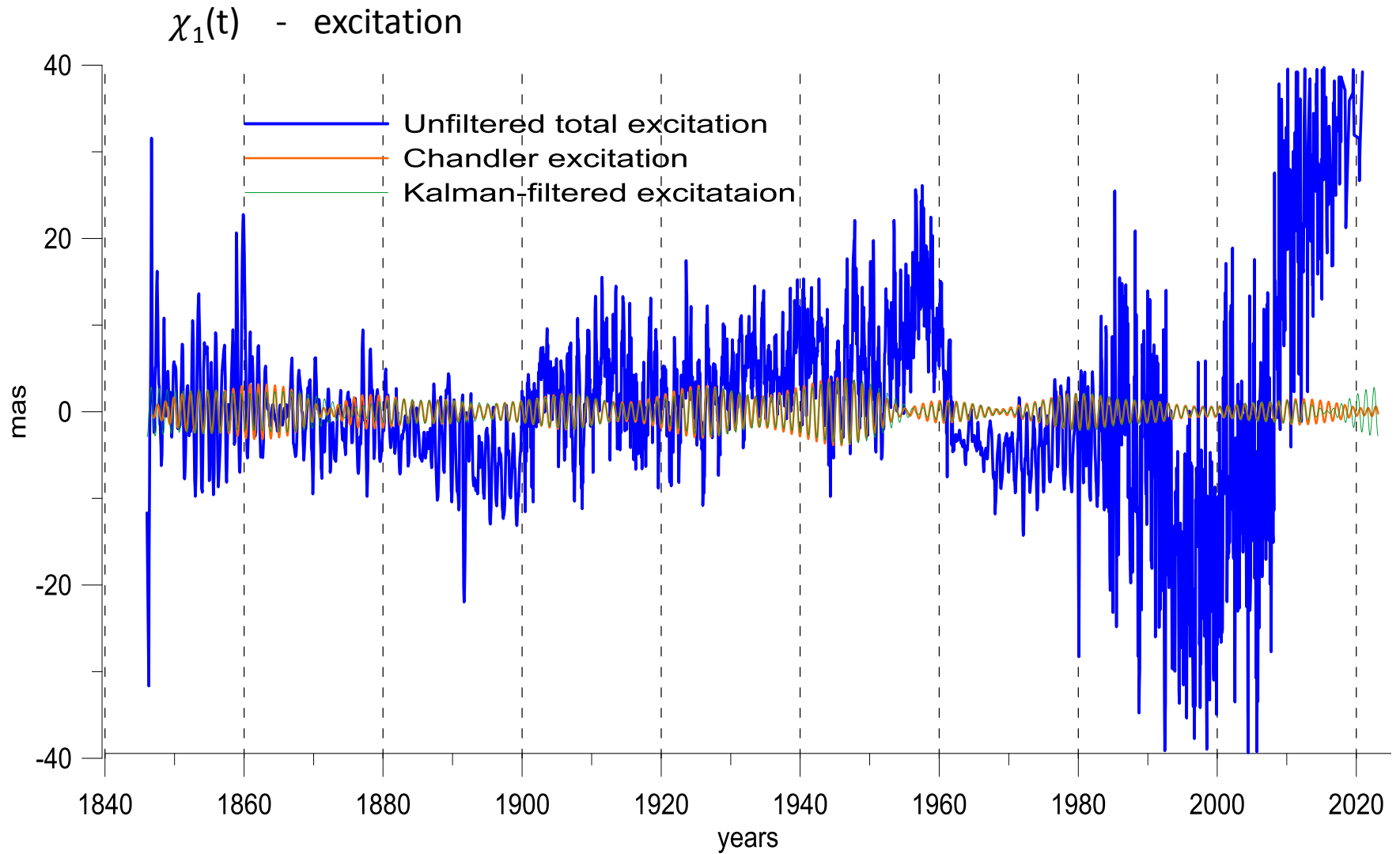
It can be corrected through regularization

$$W_{corr}(f)W^{-1}(f) = \frac{W^*(f)}{W^*(f)W(f) + \alpha I},$$

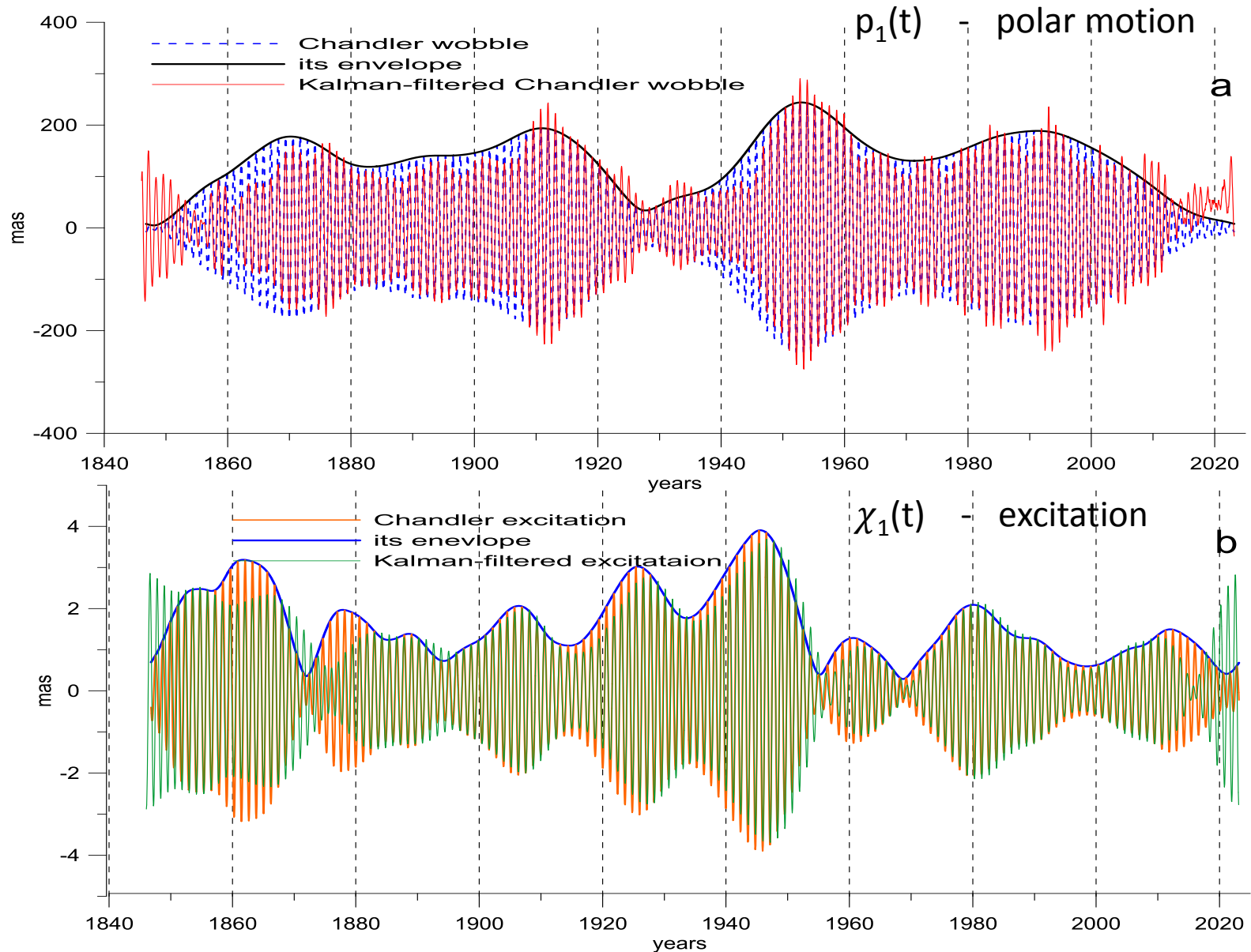
Kalman-Filtered Chandler wobble



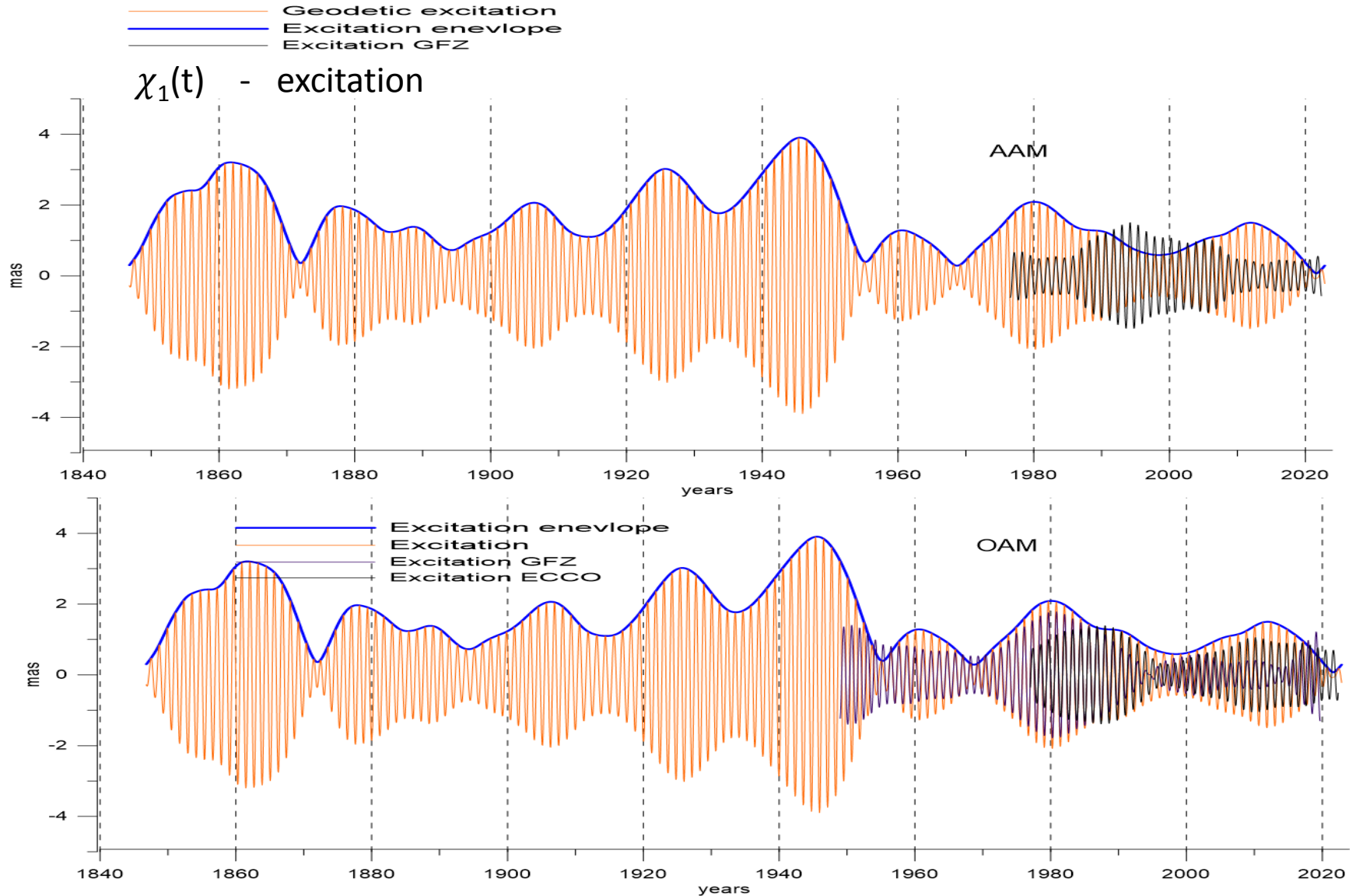
Unfiltered total excitation



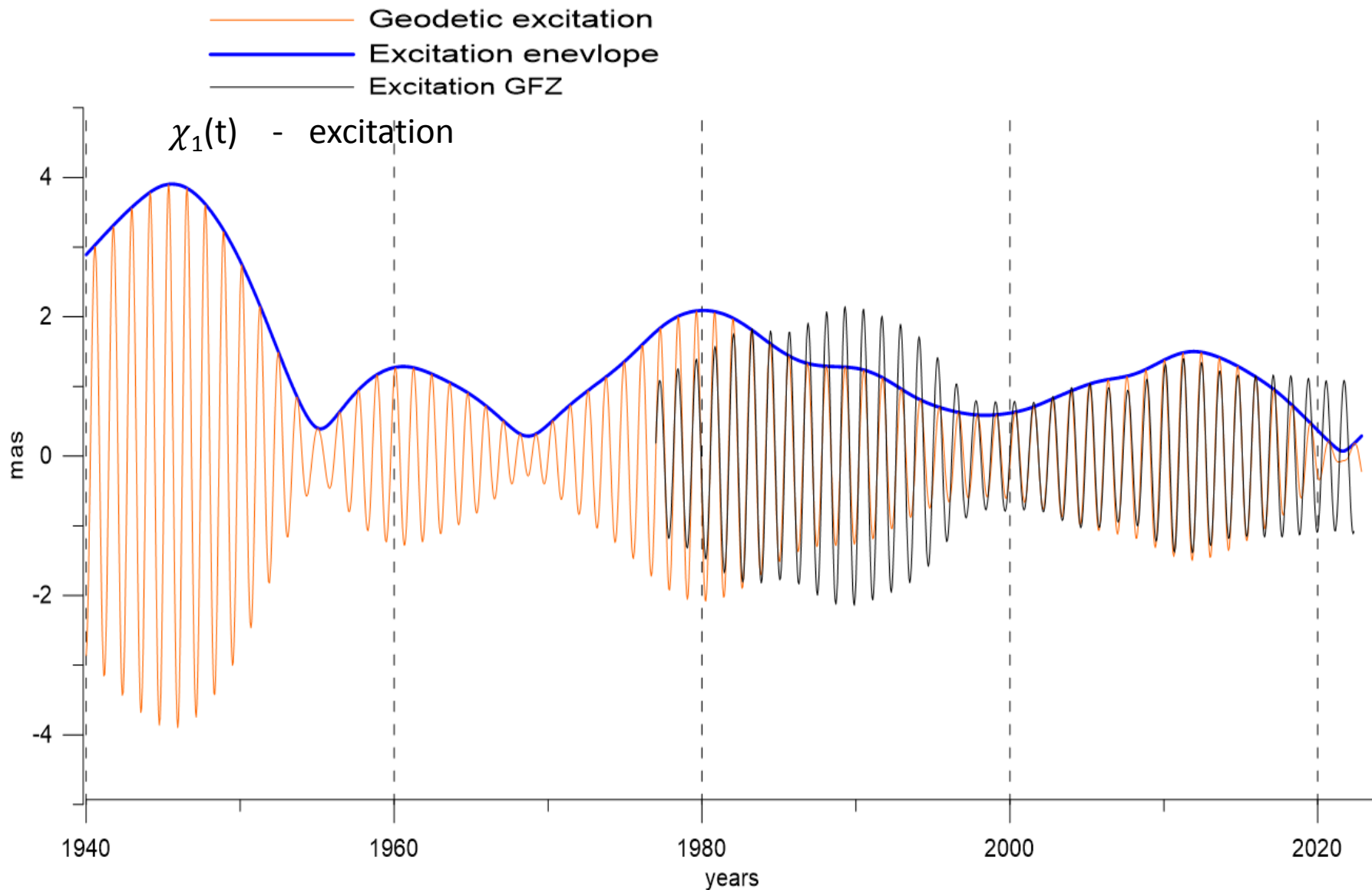
KF - Chandler wobble and its filtered excitation



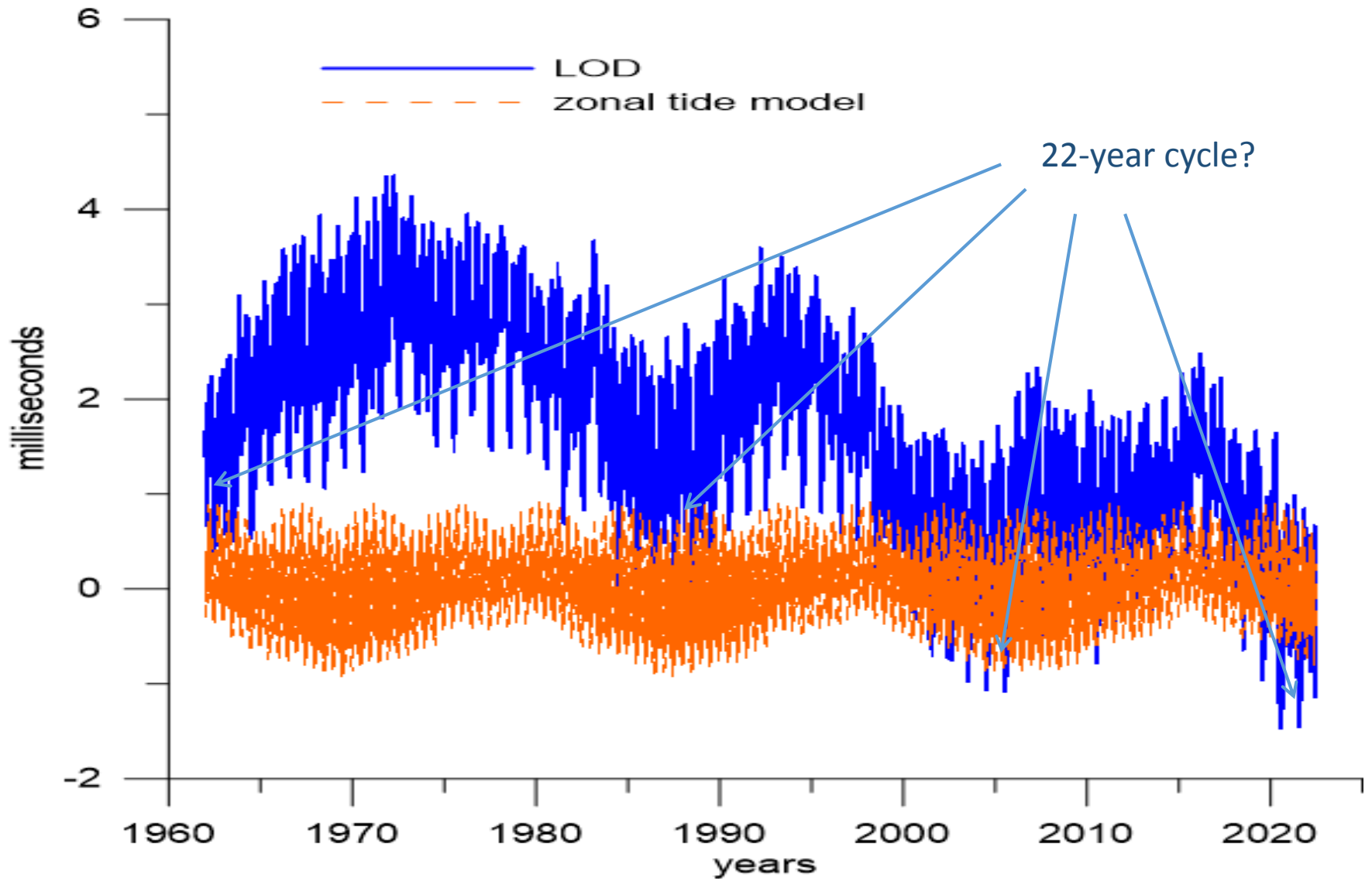
Chandler excitation compared with atmospheric and oceanic OAM and AAM (GFZ)

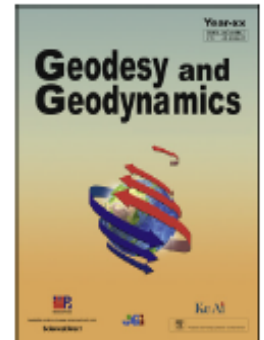


Comparison of geodetic and geophysical AAM+OAM



Length of day (LOD) from EOP C04 and zonal tides





A possible interrelation between Earth rotation and climatic variability at decadal time-scale

Leonid Zotov^{a,b,*}, C. Bizouard^c, C.K. Shum^{d,e}

^a National Research University Higher School of Economics, Moscow Institute of Electronics and Mathematics, Moscow, Russia

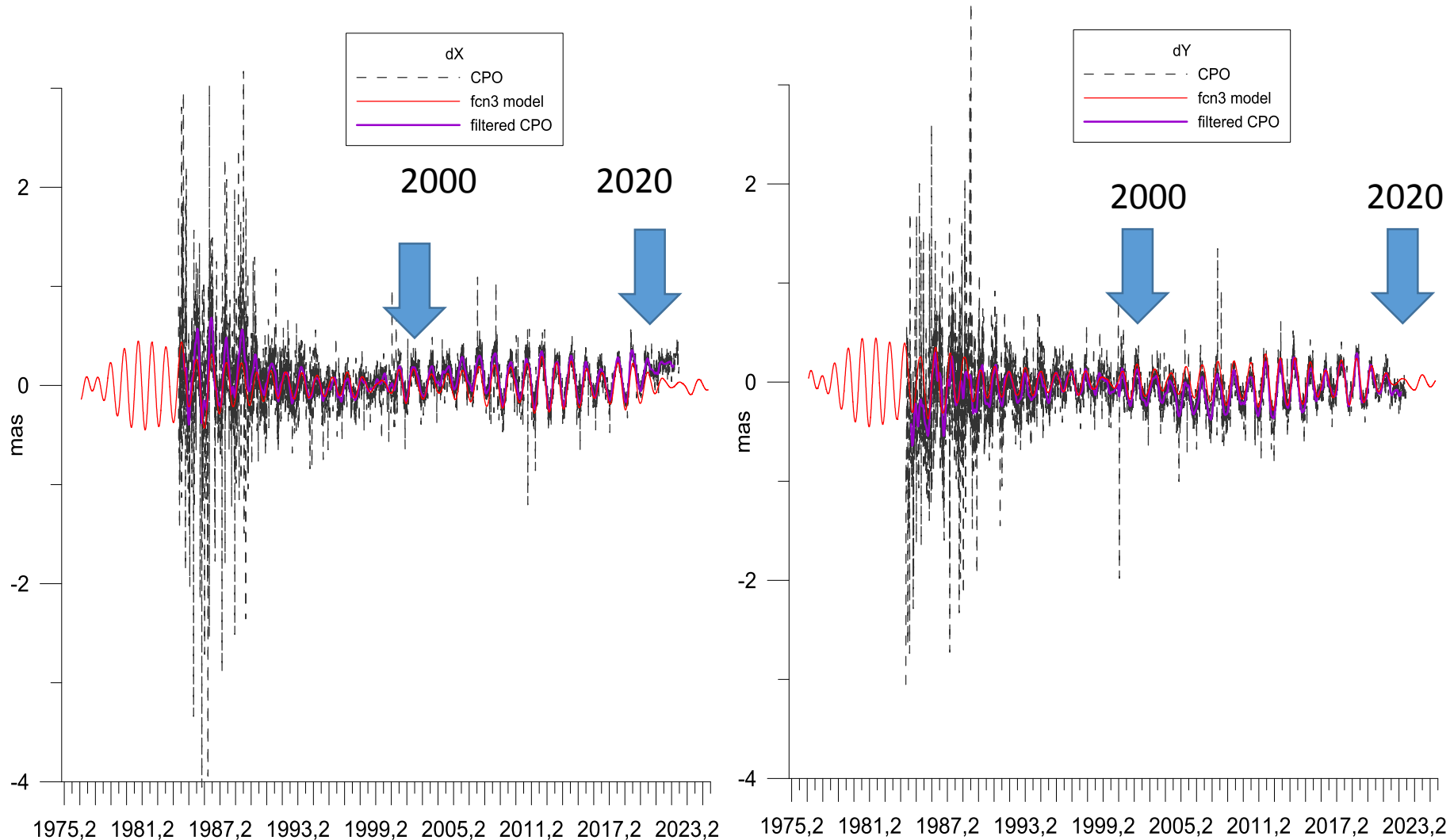
^b Lomonosov Moscow State University, Sternberg Astronomical Institute, Moscow, Russia

^c SYRTE, Observatoire de Paris, PSL Research University, CNRS, Sorbonne Universités, UPMC Univ. Paris 06, 61 avenue de l'Observatoire, 75014 Paris, France

^d Division of Geodetic Science, School of Earth Sciences, The Ohio State University, USA

^e State Key Laboratory of Geodesy and Earth's Dynamics, Institute of Geodesy & Geophysics, Chinese Academy of Sciences, Wuhan, China

Free core nutation disappeared?



Big new moon

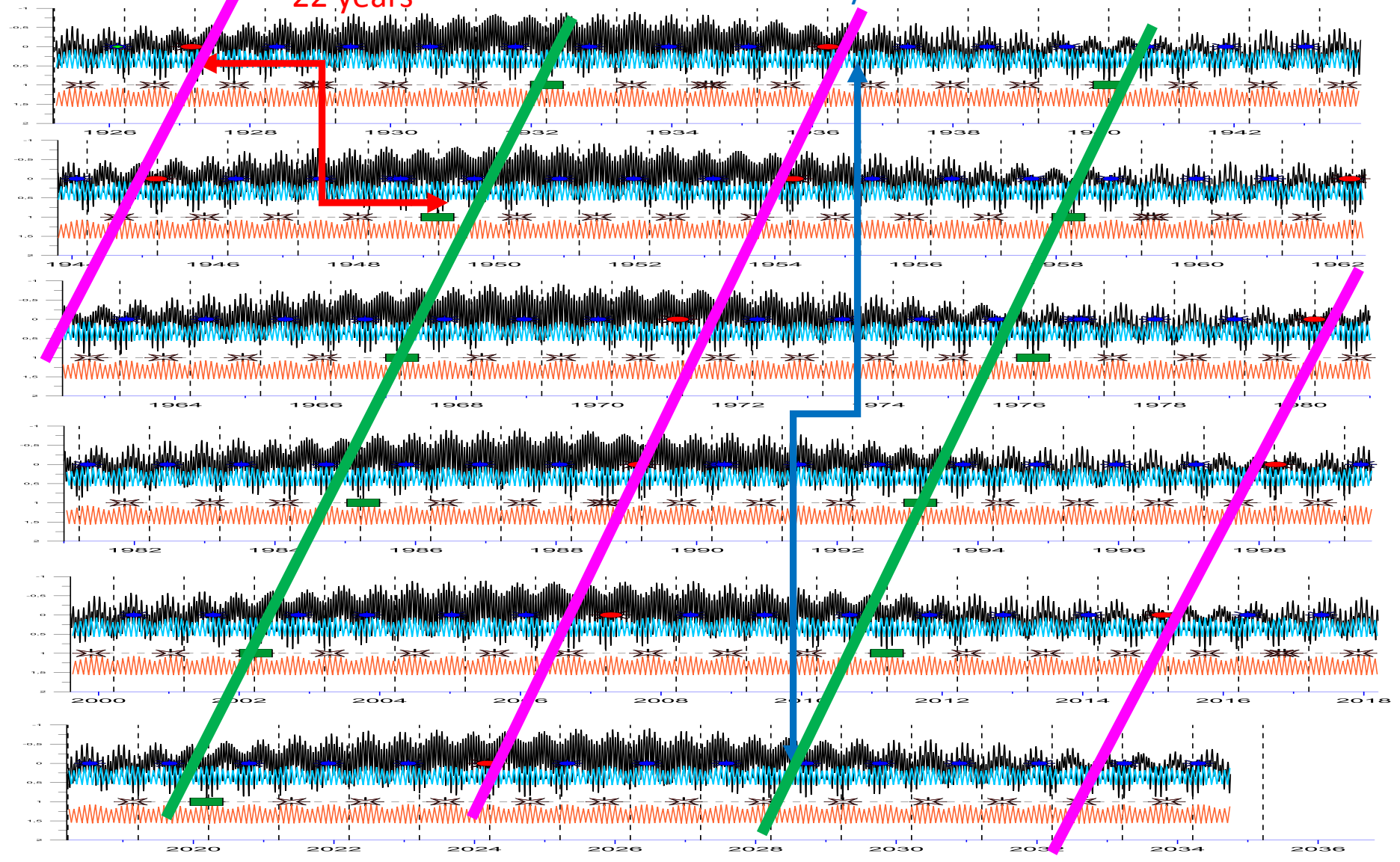
in φ

Big full moon

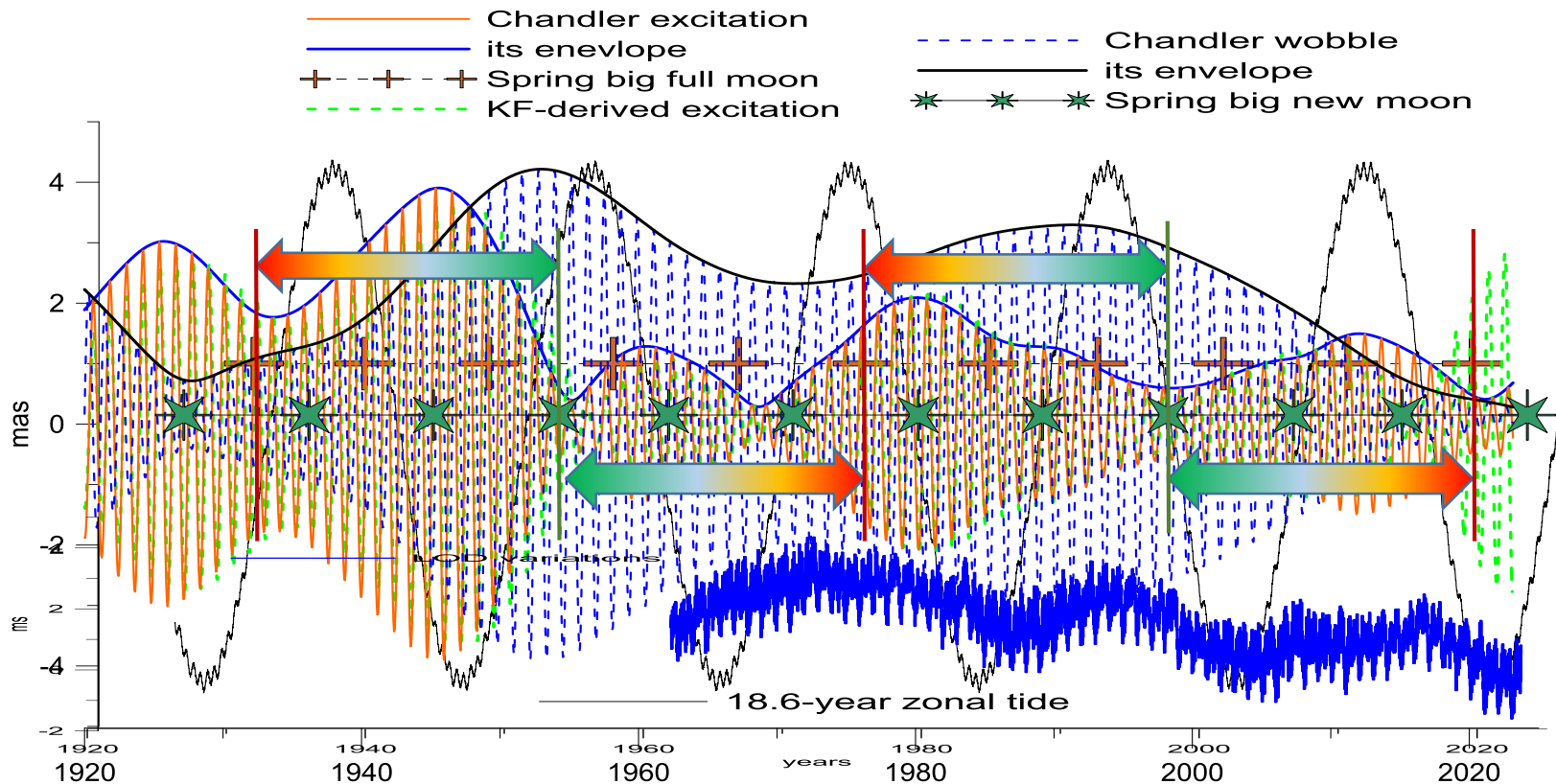
The Big moon (full/new moon in perigee) drift over precession cycle explains 90-year and 22-year modulations in the Chandler wobble

~ 22 years

~ 90 years



Chandler wobble modulations can be explained by the Moon

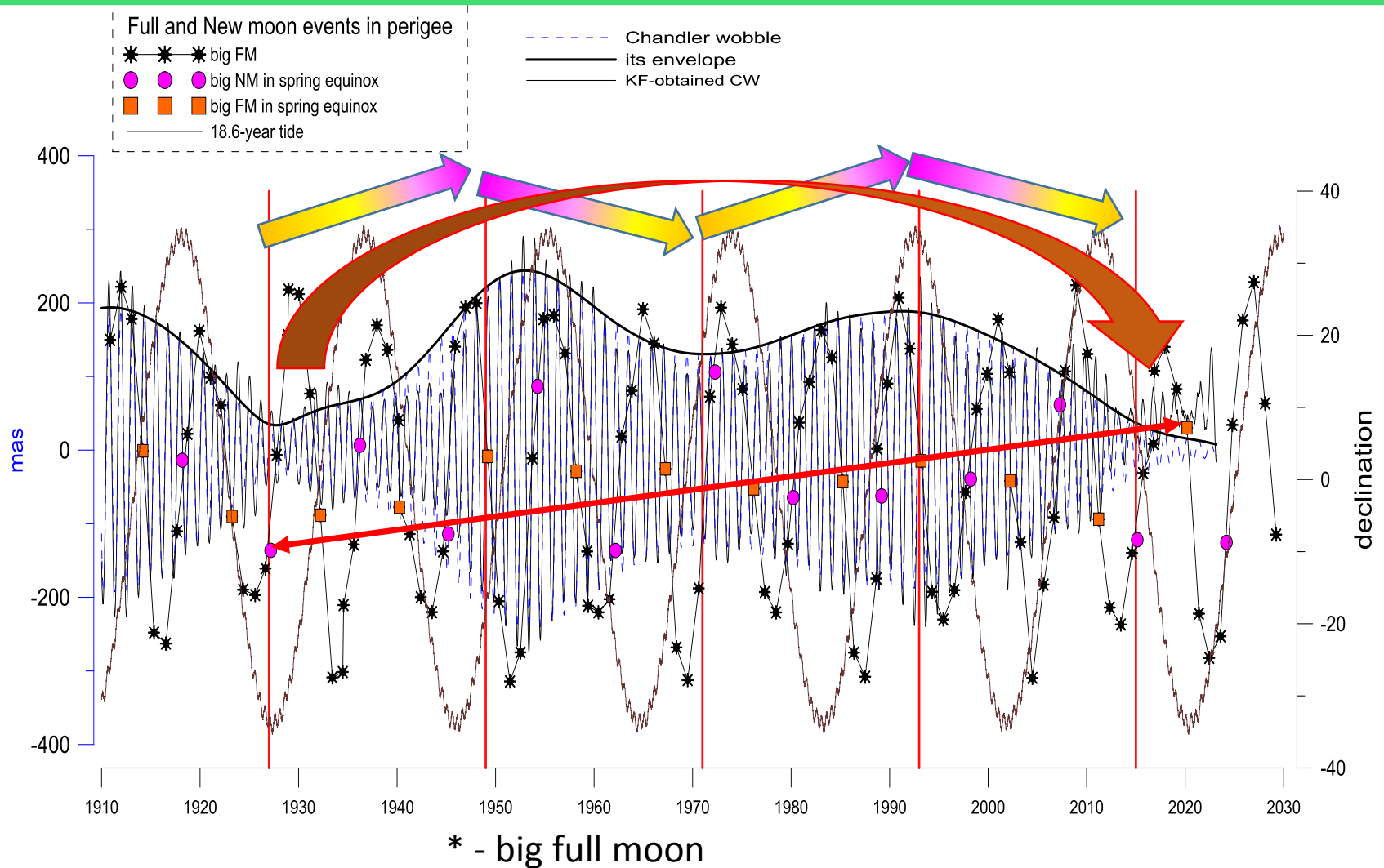


$$\frac{1}{T_{chandler}} = \frac{1}{412} - \frac{1}{22 * 365} = \frac{1}{433}$$

Zotov L.V. et al., ICCC workshop, 29 March 2023

Ideas about big moon from Yu Avsyuk and tidal cycle from N. Sidorenkov

Chandler wobble is excited by the series of syzygies occurring at different declinations



Conclusions

- We tried to develop regularized Kalman filter procedure to extract Chandler wobble, we see that it disappeared in 2017-2020 and now appeared again, its phase is now changing like in the 1930-s
- Atmospheric angular momentum AAM, oceanic angular momentum OAM, hydrological angular momentum and Sea Level changes can not quite explain acceleration of the Earth since 2016 and Chandler wobble disappearing in 2020-s
- If it is a mass term, we would see it in the gravity field
- Decadal LOD changes can be related to the oscillations in the core and/or lunar tide influence on geophysical processes
- Free Core Nutation disappeared as well
- Earth rotation velocity and Chandler wobble changes can be produced by the regular mechanism which we believe is related to the drift of the Big Moon events over the precession cycle

References

1. Zotov L.V. et al., Anomalies of the Chandler wobble of the pole in 2010-s, Moscow University Physics Bulletin, N3, p. 1-12, 2022
2. Zotov L.V., Earth rotation and climate processes, monography, Moscow, MIEM HSE, 2022
3. Sidorenkov N.S. Synchronization of terrestrial processes with frequencies of the Earth-Moon-Sun system (AApTr), 2017, Vol. 30, Issue 2, pp. 249-260.
4. Malkin Z., Belda S., Modiri S. Detection of a New Large Free Core Nutation Phase Jump, Sensors 2022, 22, 5960, <https://doi.org/10.3390/s22165960>
5. Bizouard Ch., Geophysical modelling of the polar motion, De Gruyter, 2020.
6. Sidorenkov N.S., The Interaction Between Earth's Rotation and Geophysical Processes, Wiley-VCH Verlag, Weinheim, 2009
7. Avsuk, Yu.N. and Maslov, L.A., Long period tidal force variations and regularities in orbital motion of the Earth–Moon binary planet system, Earth, Moon Planets, 2011, vol. 108, no. 1, pp. 77–85.
8. Fedorov, E. P. & Yatskiv, Y. S., The Cause of the Apparent "Bifurcation" of the Free Nutation Period. Soviet Astronomy, Vol. 8, p.608, 1965.
9. Malkin Z., Miller N. Chandler wobble: two more large phase jumps revealed, Earth Planets Space, 62, 943–947, 2010.
10. Wai Yan Soe, Study of the parameters of Earth's polar motion dependence on the Lunar orbit precession, PhD thesis, Moscow, 2022
11. Graham Jones, Earth Sets New Record for Shortest day. Time&Date, 27.07.22 <https://www.timeanddate.com/news/astronomy/shortest-day-2022>

Thank you for attention!

